

Plasma physics

2023 winter s. 2/1 , C+Ex

prof. Juraj Glosík, doc. Radek Plašil

- Negative ions in plasma
- Elementary processes: recombination



Negative ions—anions—in plasma

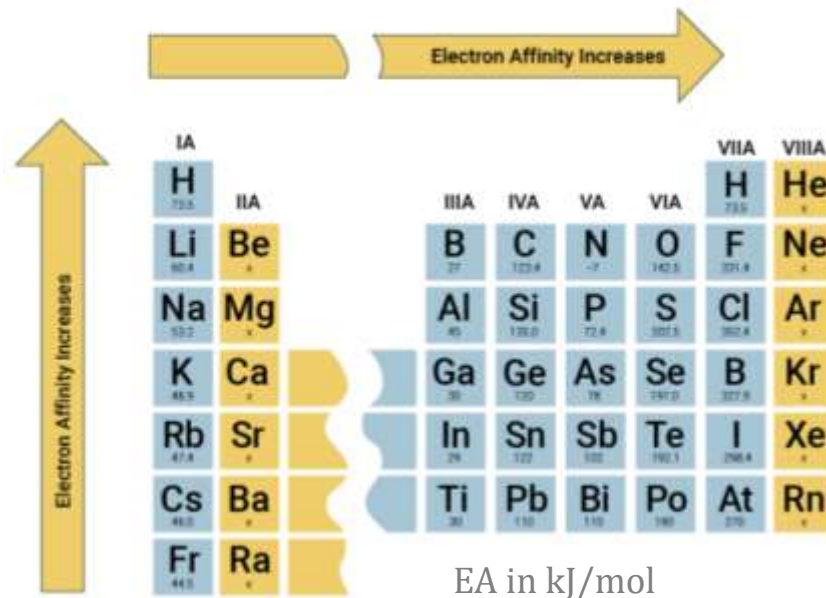
Electronegative plasma
boundary condition for
validity of Bohm criterion

$$\frac{T_e}{T_{\text{anion}}} \approx 10$$

$$\frac{n_{\text{anion}}}{n_e} \approx 3$$

$$\frac{n_{\text{anion}}}{n_e} \approx \sqrt{\frac{T_e}{T_{\text{anion}}}}$$

$$\frac{n_{\text{anion}}}{n_e} \approx 10$$



Species	Electron Affinity “EA”
O ₂	0.45 eV
H	0.75 eV
C	1.3 eV
O	1.5 eV
Cl	3.6 eV
C ₆ H	3.8 eV

McCarthy M. C. et al. *ApJ* 652 (2006) L141
doi: 10.1086/510238

"Electron Affinity of The Elements" Published on Jun 19, 2021.

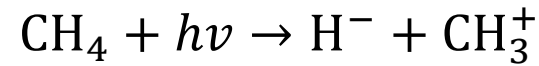
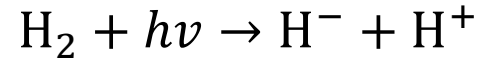
<https://breakingatom.com/learn-the-periodic-table/electron-affinity-of-the-elements>

<https://webbook.nist.gov/chemistry/ea-ser/>



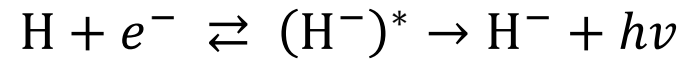
Formation of anions in plasma

- Ion-pair formation



Maximum cross sections
 $\approx 20 \text{ eV}$

- Radiative electron attachment



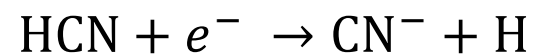
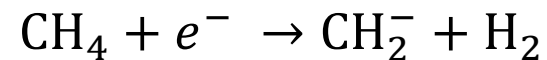
$$k = 3 \cdot 10^{-16} \text{ cm}^3 \text{ s}^{-1}$$

at 300 K



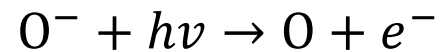
$$k = 6 \cdot 10^{-8} \text{ cm}^3 \text{ s}^{-1}$$

- Dissociative attachment

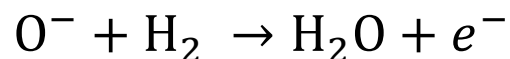
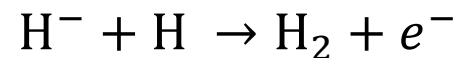


Destruction of anions

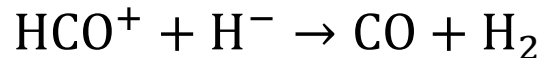
- Photodetachment



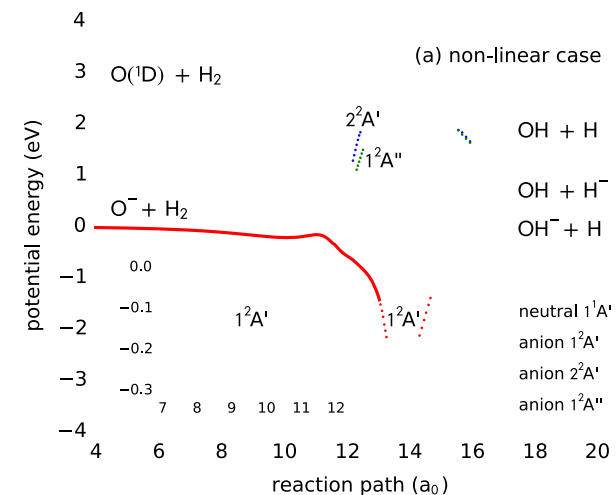
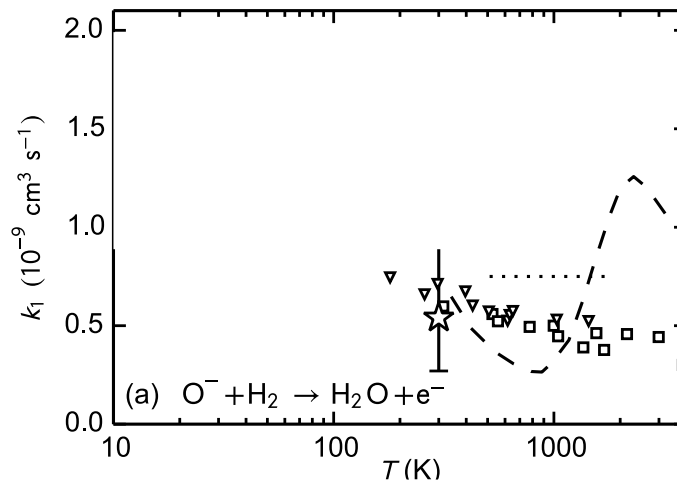
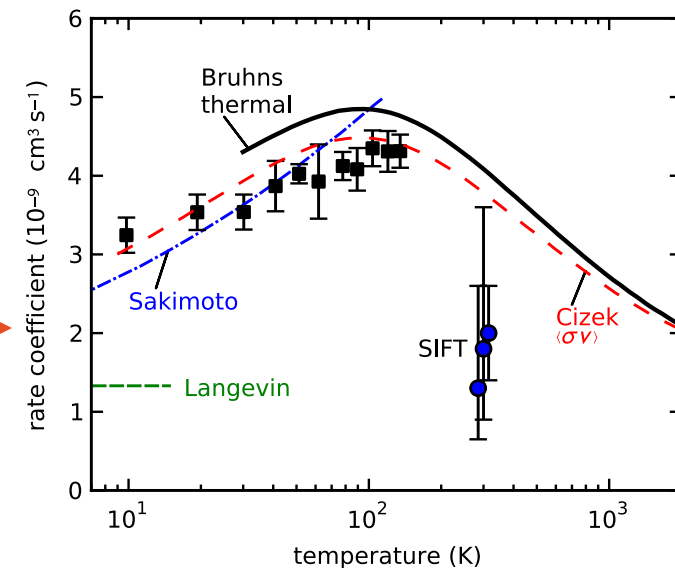
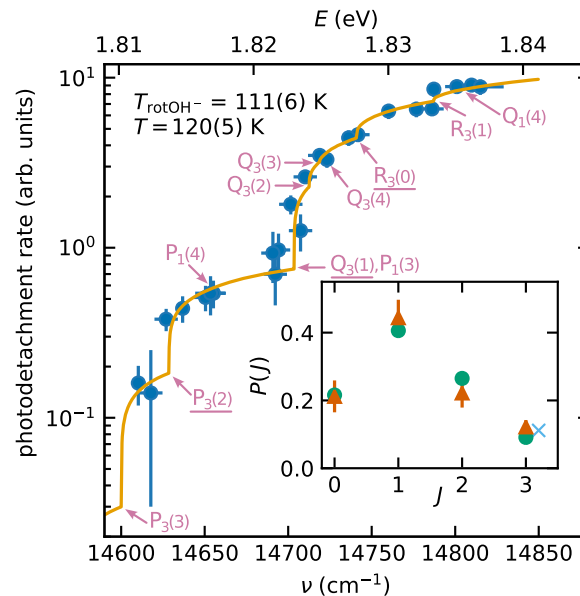
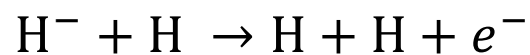
- Associative detachment



- Cation-Anion recombination



- Electron detachment in collision



Solar opacity

Transitions

- bond-bond transitions
 $h\nu + H(n_1) \rightarrow H(n_2 > n_1)$
- free-free transitions
 inverse Bremsstrahlung
 $h\nu_1 + e^- + H^+ \rightarrow h\nu_2 + e^- + H^+$
- bond-free transition
 $h\nu + H(n) \rightarrow H^+ + e^-$

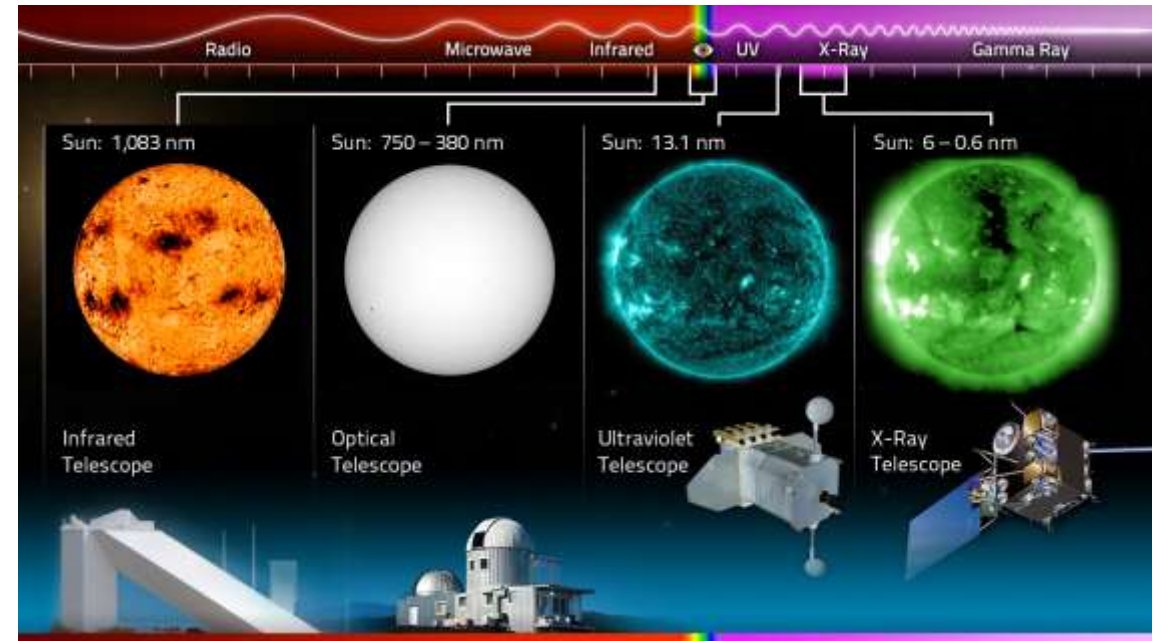
$$n = 1 \Rightarrow 13.6 \text{ eV}; n = 2 \Rightarrow 3.4 \text{ eV}$$



$$n = 1 \Rightarrow 0.75 \text{ eV}$$

electron affinity

Photosphere of Sun 6500 to 4200 K
 the energy is transferred by radiation, convection is stopped
 photon mean free path 1 to 100 km



Multiwavelength Astronomy project
ecuip.lib.uchicago.edu/multiwavelength-astronomy/astrophysics/02.html

visible 1.6 – 3.3 eV

Sun

$$\frac{[H^-]}{[H]} \approx 10^{-7}$$

$$\frac{[H(n \geq 3)]}{[H]} \approx 10^{-9}$$

The continuum radiation is mostly formed by bond-free transitions of the anion H^- .
 It is very similar to the black body radiation with a maximum at $\lambda = 500 \text{ nm}$.



Recombination of positive ions

Recombination rate coefficient $\alpha \equiv k_{\text{recombination}}$

Net recombination loss due to one type of cation "i"—for example O_2^+

$$\zeta(t) = \frac{n_i}{n_e}$$

$$\frac{dn_e}{dt} = -\alpha n_i n_e = -\alpha \zeta(t) n_e^2$$

Afterglow plasma Recombination and Diffusion losses

In case of one dominant ion species

$$\zeta(t) = 1$$

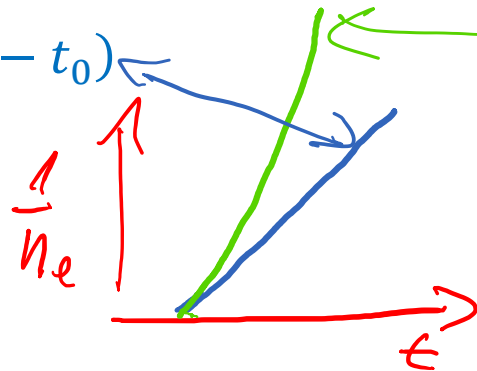
$$\frac{dn_e}{dt} = -\alpha n_e^2 - D_a \nabla^2 n_e$$

$$D_a \nabla^2 n_e = \frac{D_a}{\Lambda^2} n_e = \frac{1}{\tau_D} n_e$$

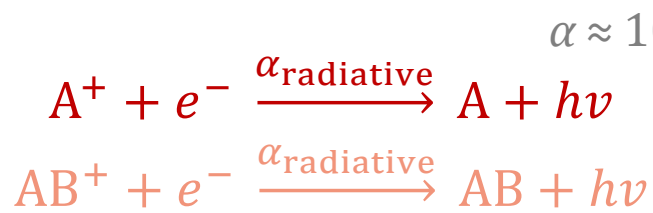
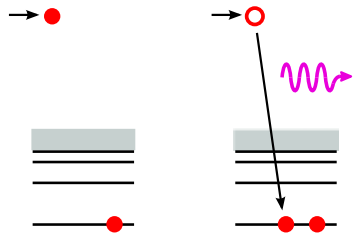
$$\frac{dn_e}{dt} = -\alpha n_e^2$$

$$\frac{dn_e}{dt} = -\alpha n_e^2 - \frac{n_e}{\tau_D}$$

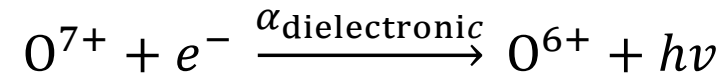
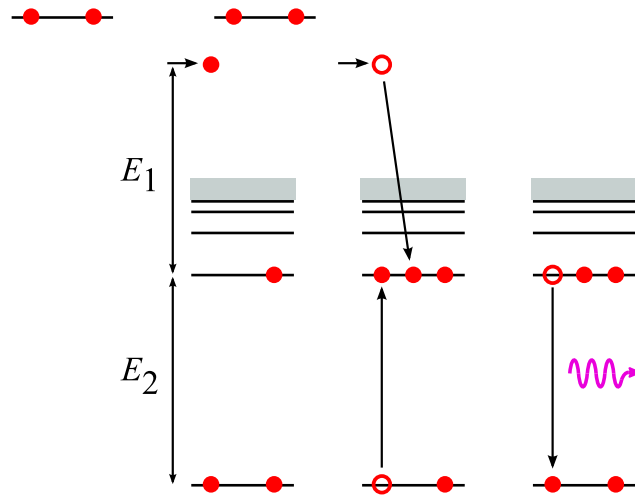
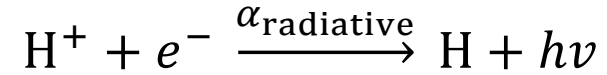
$$\frac{1}{n_e(t)} = \frac{1}{n_e(t_0)} + \alpha(t - t_0) \quad \frac{1}{n_e(t)} = \frac{1}{n_e(t_0)} \exp\left(\frac{t}{\tau_D}\right) + \alpha \tau_D \left[\exp\left(\frac{t}{\tau_D}\right) - 1 \right]$$



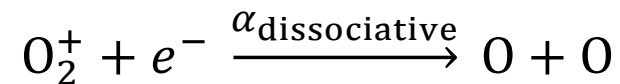
Recombination of positive ions



$$\alpha \approx 10^{-10} \text{ cm}^3 \text{ s}^{-1}$$

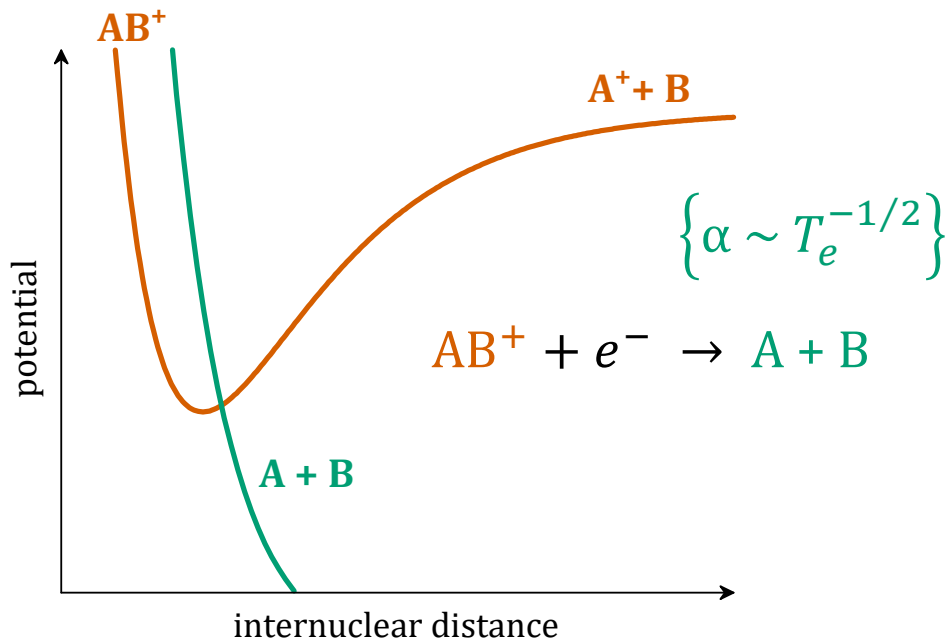


$$\alpha \approx 10^{-10} - 10^{-7} \text{ cm}^3 \text{ s}^{-1}$$

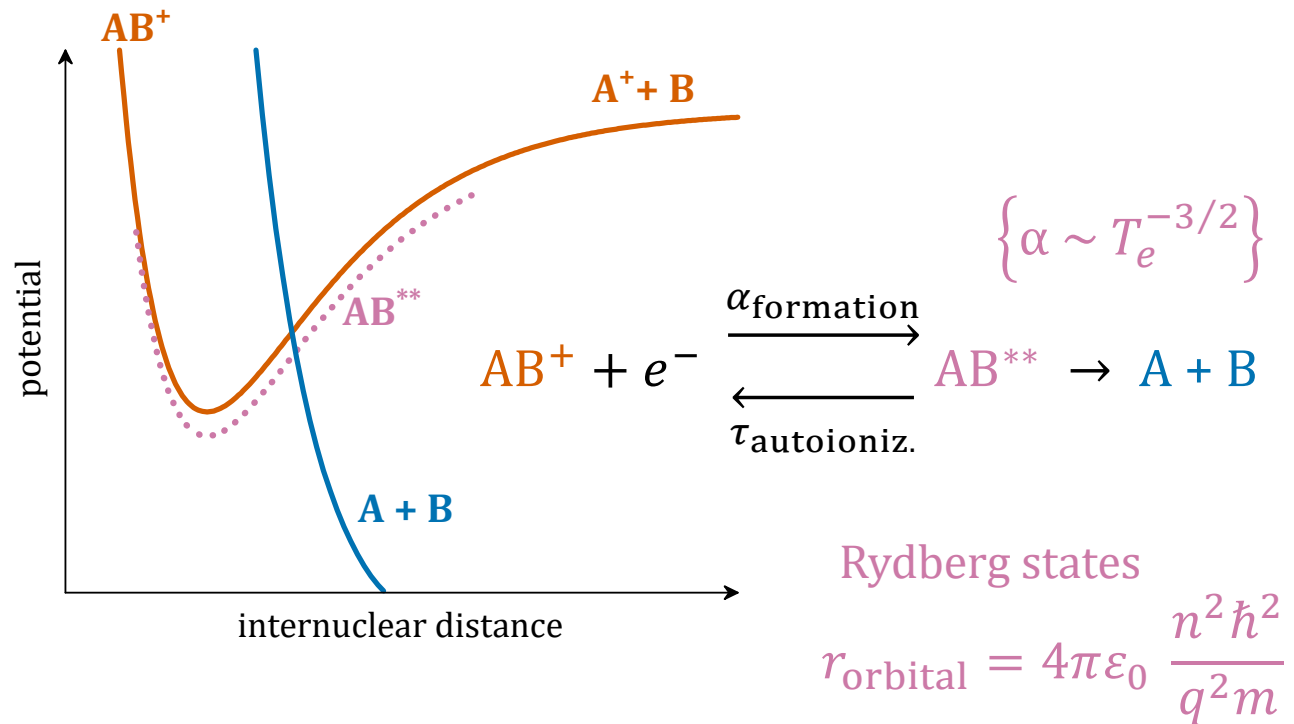


Dissociative recombination of positive ions

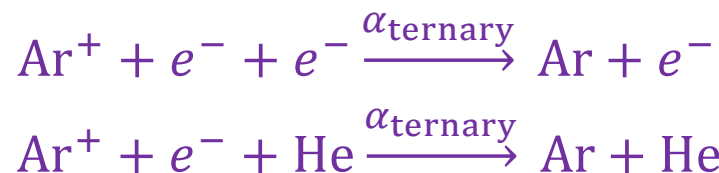
Direct recombination



Indirect recombination



Bardsley J. N., Biondi M. A. 1970. Advances in atomic and molecular physics.
In: DR Bates, editors. New York: Academic Press. pp 1-57.



Collisional Radiative Recombination

Neutral Assisted Recombination



Example of dissociative recombination O_2^+

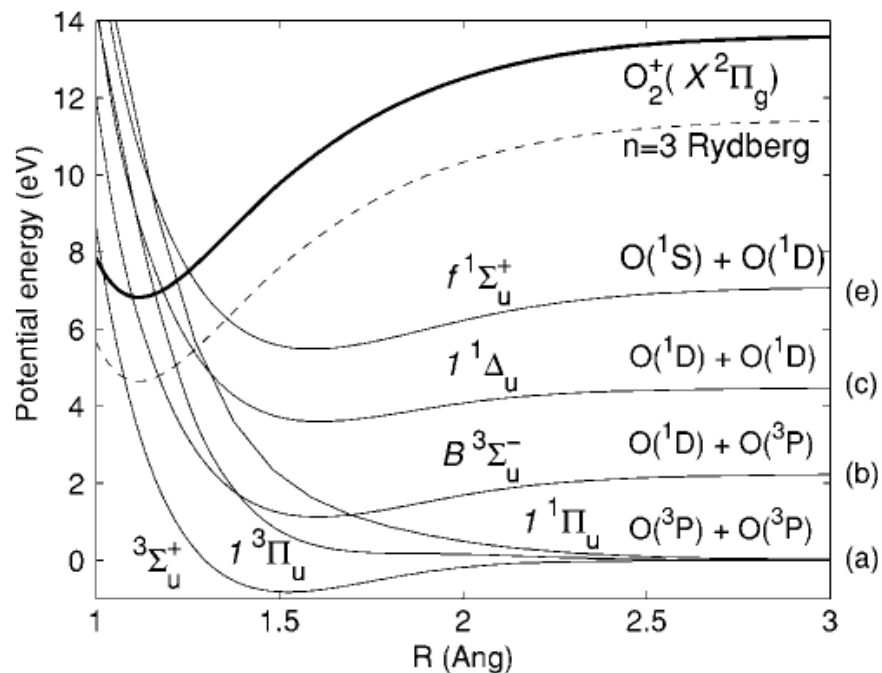


FIG. 1. Schematic of the diabatic potential curves relevant for the DR of O_2^+ . The dissociation limits connected with each valence capture state are given on the right. The labels (a)–(c), and (e) refer to Eqs. (1a)–(1c) and (1e), respectively.

Petrignani A. et al. *J. Chem. Phys.* 122 (2005) 23431
doi: 10.1063/1.1937388.

Peverall R. et al. *J. Chem. Phys.* 114 (2001) 6679
doi: 10.1063/1.1349079

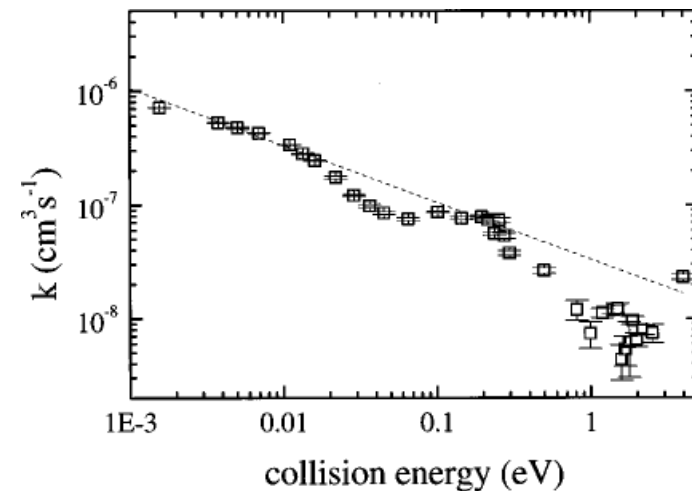
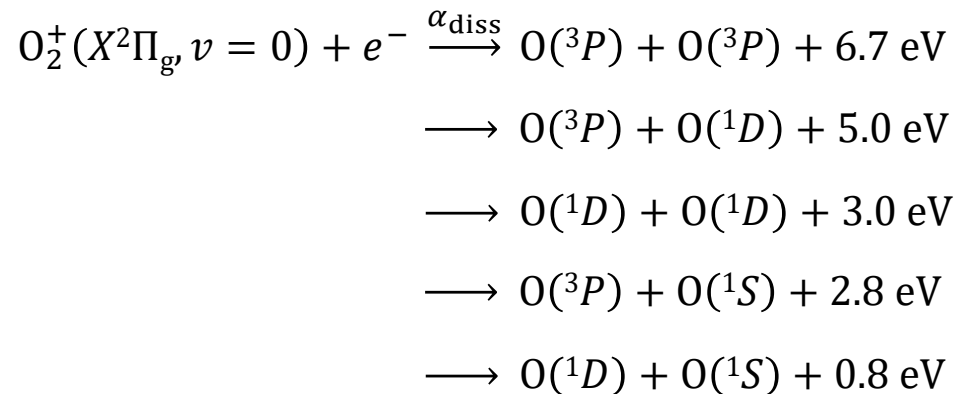


FIG. 2. DR rate coefficient k as a function of electron collision energy from 1 meV to 5 eV. Statistical errors are shown at the 1σ level. The dotted line shows the threshold $E^{-1/2}$ behavior. Both the rate coefficient and the energy are shown on a logarithmic scale.



DR without curve crossing

Tunneling dissociative recombination

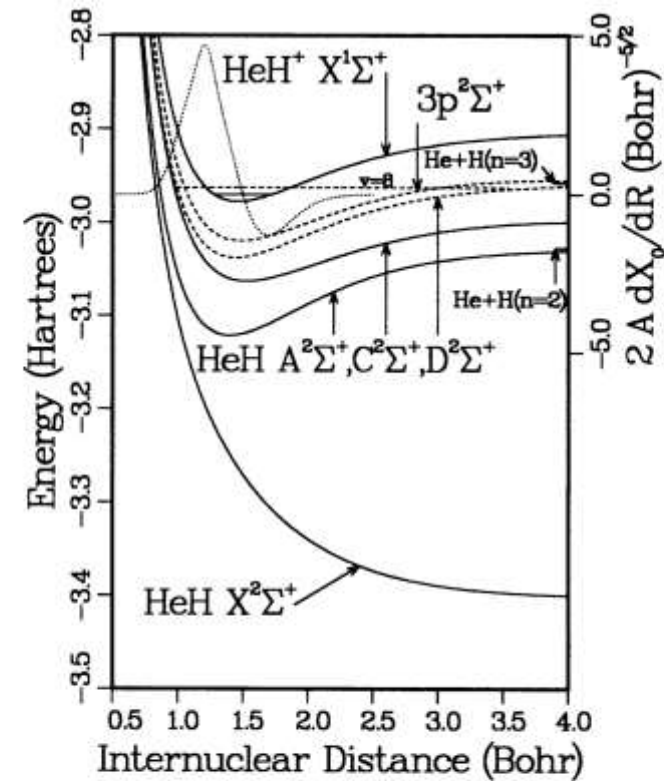
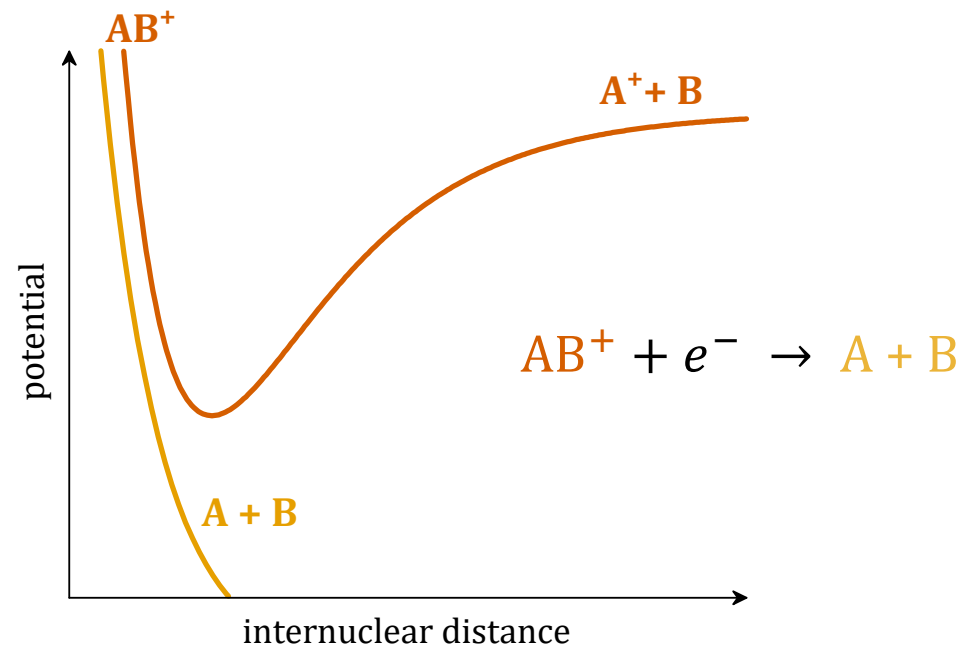


FIG. 1. The calculated HeH dissociative potential-energy curves (solid lines), the $n = 3$ Rydberg states (dashed lines), and the HeH^+ ground state (solid line) with the $v = 0$ level are shown. The $v = 8$ resonance level of the $D^2\Sigma^+$ state is shown as the dashed line. The dotted line shows the product, $2A(R)d\chi_0/dR$, with its ordinate axis on the right.



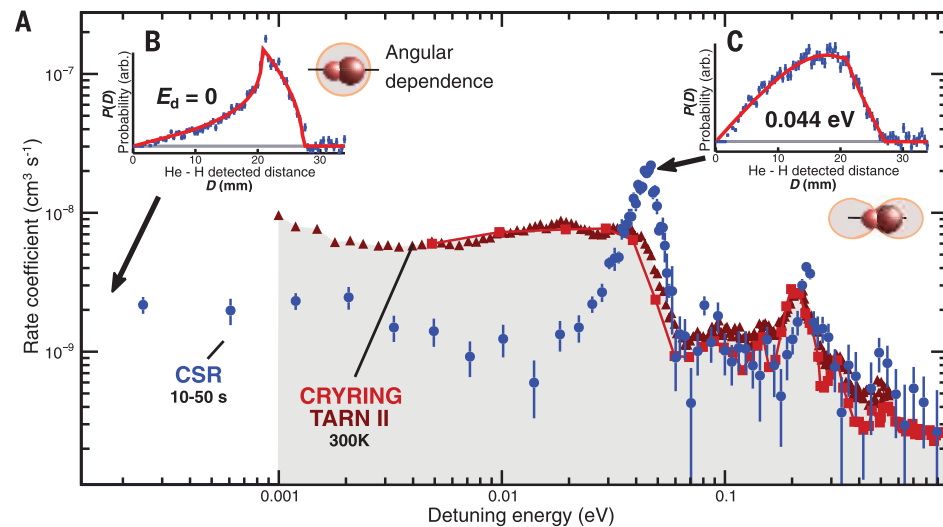
DR of HeH⁺

Astrophysical detection of the helium hydride ion HeH⁺

Rolf Güsten^{1*}, Helmut Wiesemeyer¹, David Neufeld²,
Christophe Risacher^{1,5} & Jürgen Stutzki³

Güsten R. et al. *Nature* 568 (2019) 357.
doi: 10.1038/s41586-019-1090-x

Reaction	Rate coefficient [cm ³ s ⁻¹]	Notes
He ⁺ + H → HeH ⁺ + hν	1.4 × 10 ⁻¹⁶	1
HeH ⁺ + e → He + H	3.0 × 10 ⁻¹⁰ (T/10 ⁴ K) ^{-0.47}	2
HeH ⁺ + H → He + H ₂ ⁺	1.2 × 10 ⁻⁹ (T/10 ⁴ K) ^{-0.11}	3



Novotný O. et al. *Science* 365 (2019) 676.
doi: 10.1126/science.aax5921

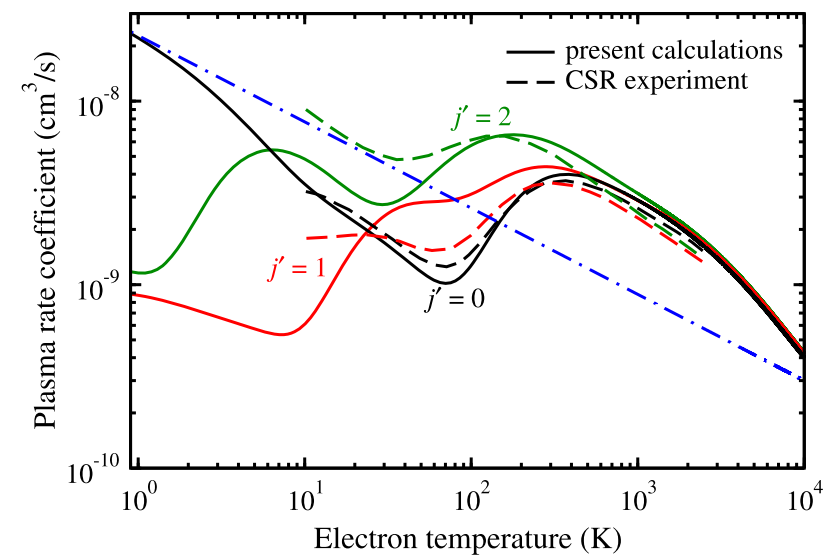


FIG. 5. Plasma rate coefficients are shown for different initial rotational states $j' = 0, 1, 2$. The solid curves result from the present theory, while the dashed curves are taken from the experiment at the CSR [4]. The dot-dashed line represents data employed by Guesten *et al.* [1] in their chemistry kinetics simulations of NGC 7027.

Čurík R. et al. *Phys. Rev. Lett.* 124 (2020) 043401
doi: 10.1103/PhysRevLett.124.043401



DR of H_3^+

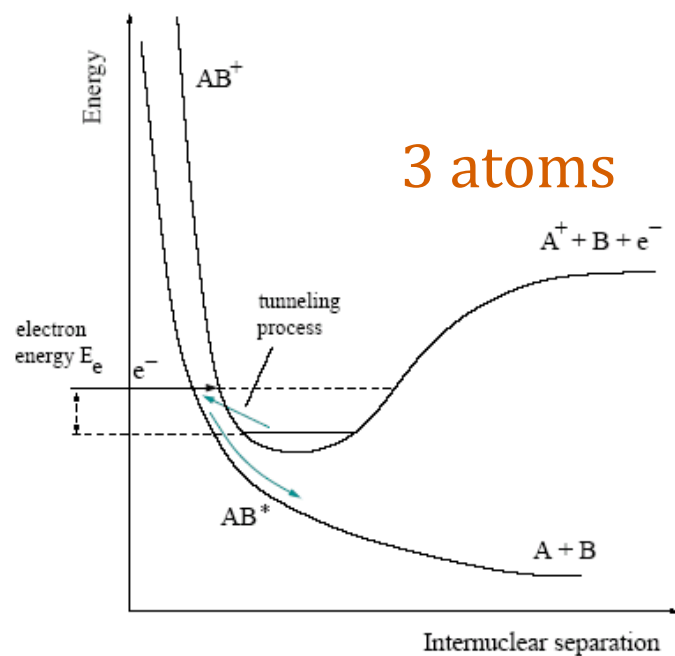
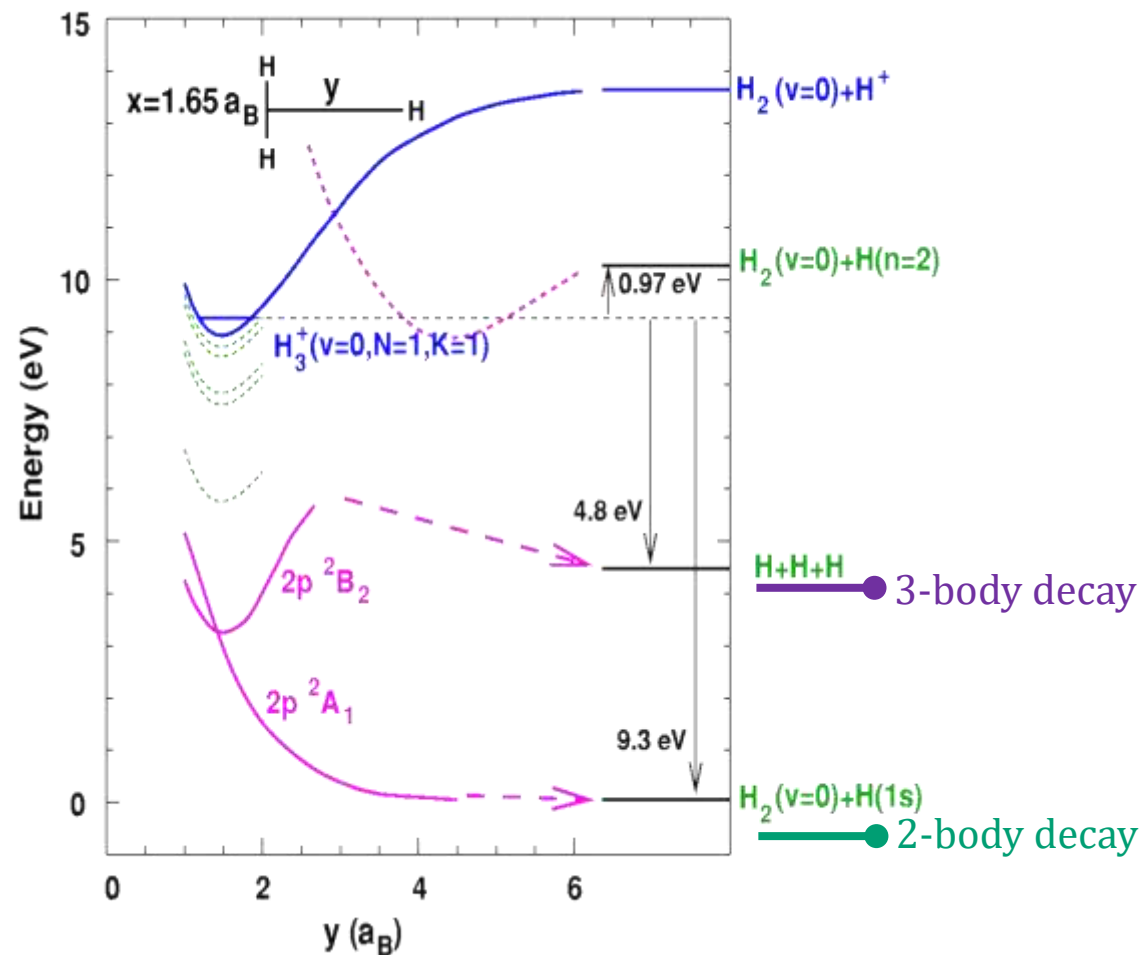


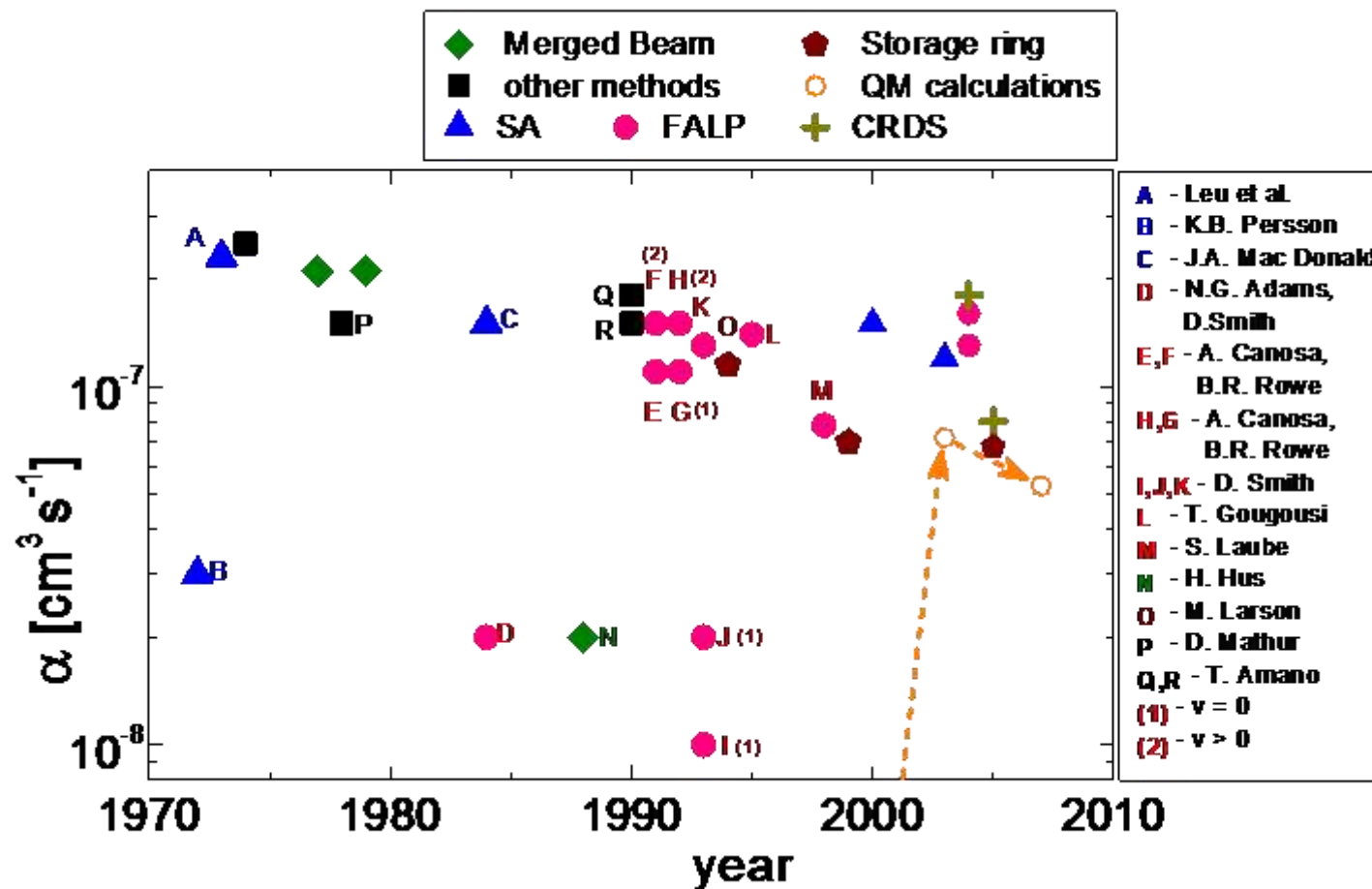
Figure 4.3: Sketch of tunneling mode dissociative recombination.

Electron capture via Jahn-Teller coupling of electronic and ro-vibrational motion

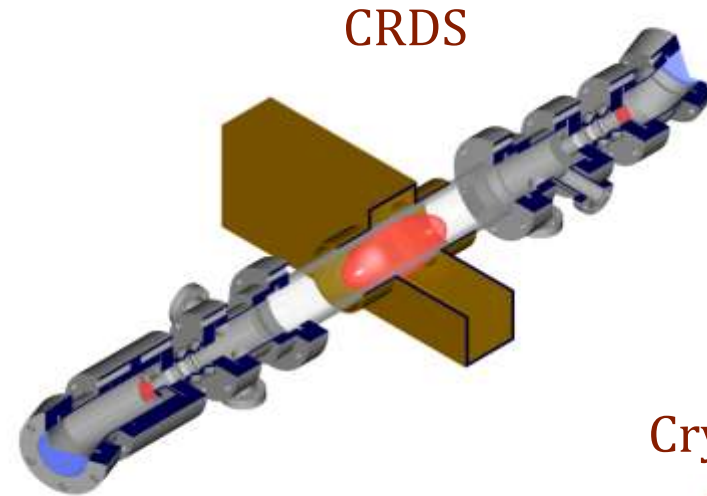
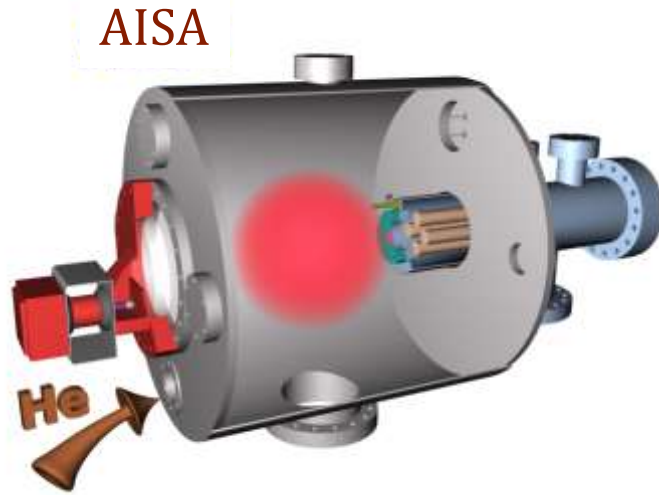


DR of H_3^+

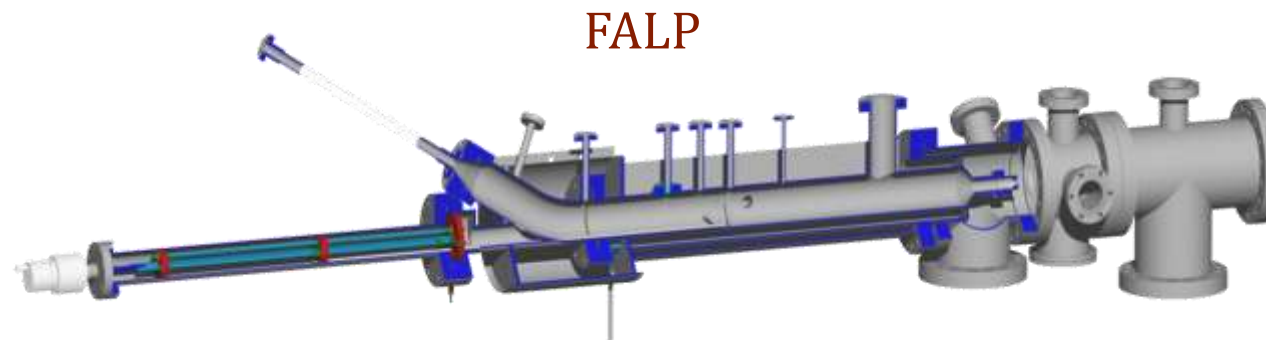
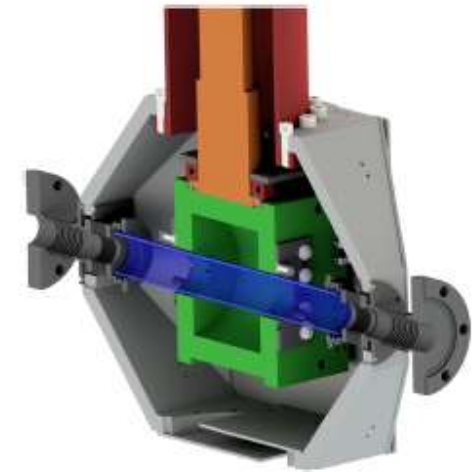
In previous years more than 25 experimental studies resulted in very different values.



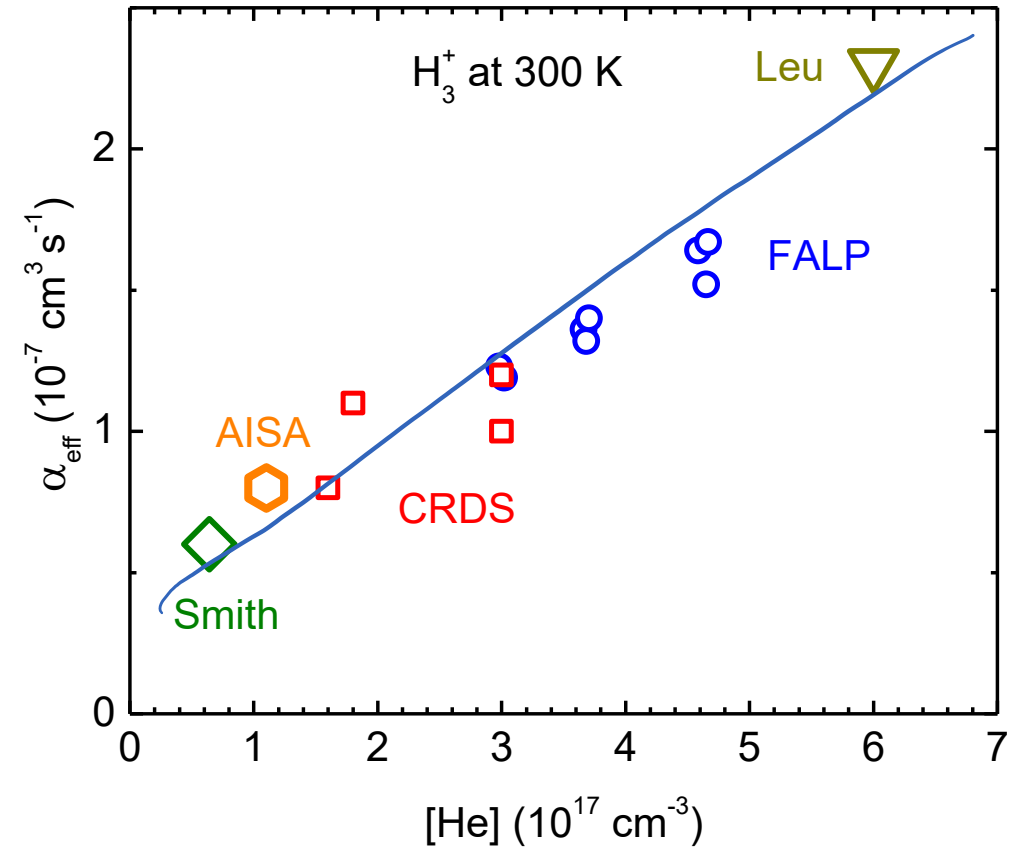
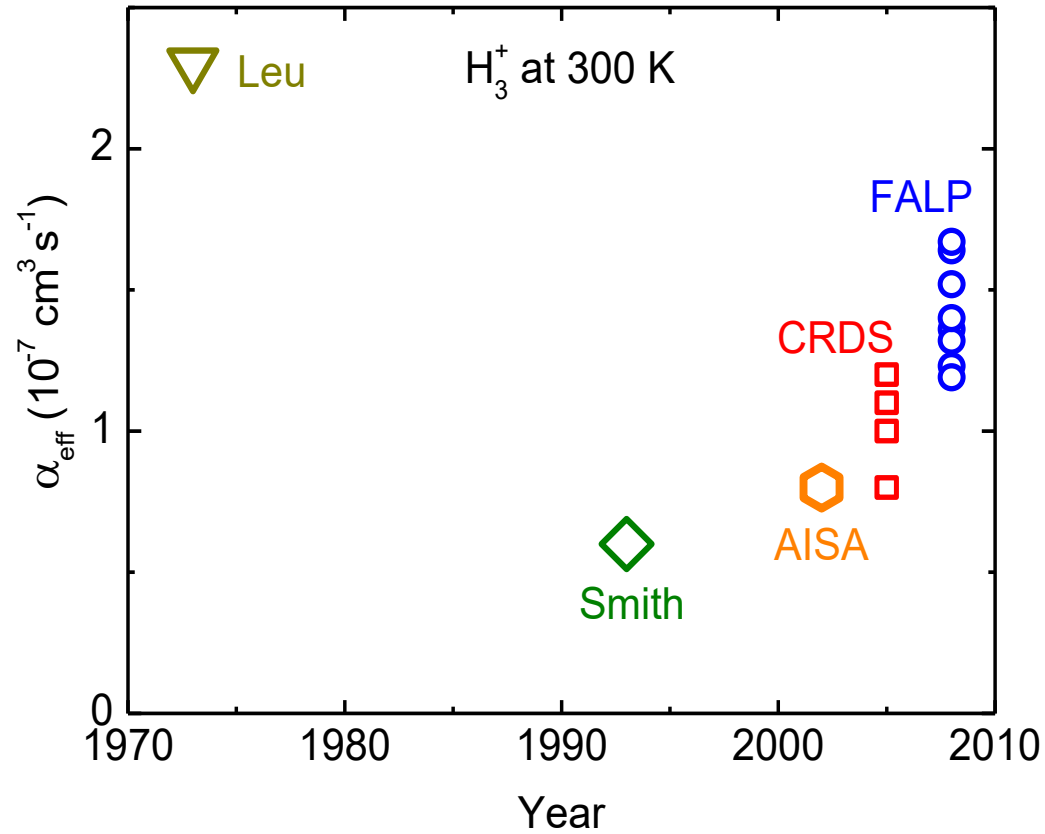
DR experiments 40 – 300 K, 100 – 10k Pa



Cryo-SA-CRDS



DR of H_3^+

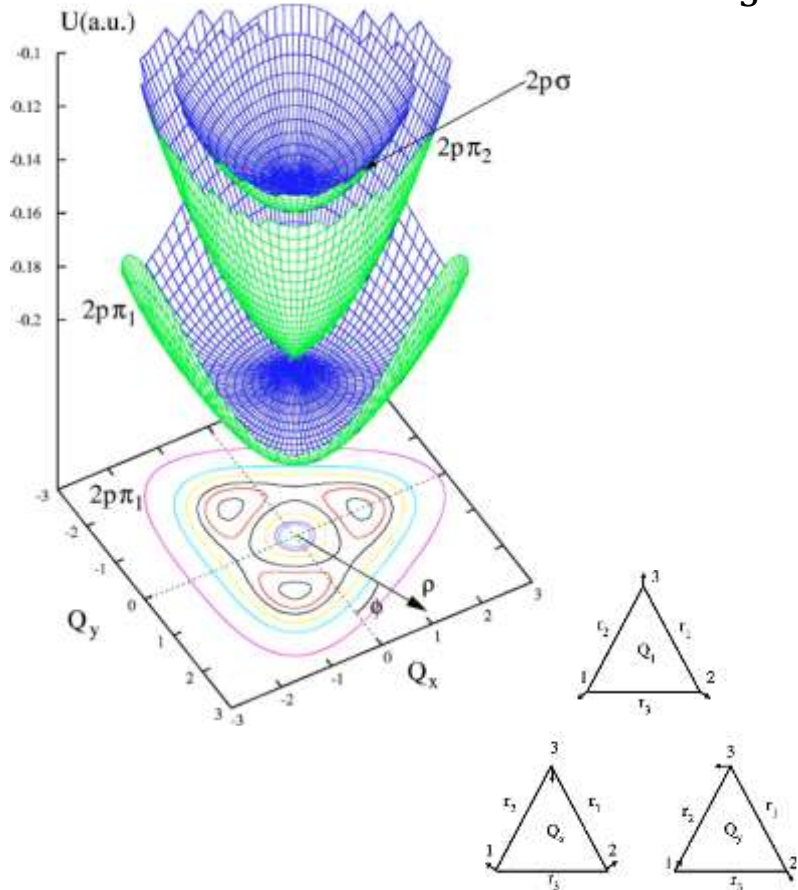
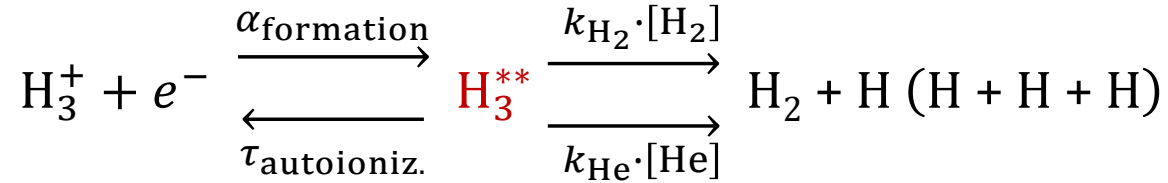


Plašil R. et al. *Int. J. Mass Spectrom.* 218 (2002) 105.
Glosík J. et al. *J. Phys.: Conf. Ser.* 4 (2005) 104.

Macko P. et al. *Int. J. Mass Spectrom.* 233 (2004) 299.
Glosík J. et al. *Phys. Rev. A* 79 (2009) 052707.

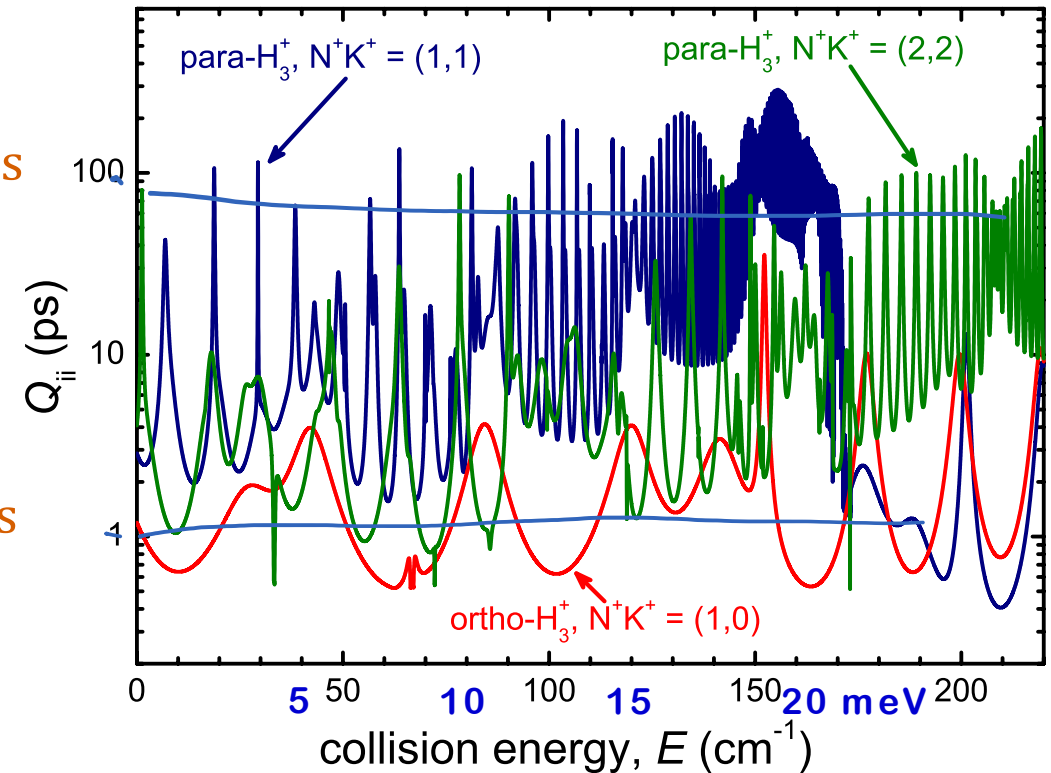


Ternary recombination of positive ions



$\tau_{\text{para}} \approx 100 \text{ ps}$

$\tau_{\text{ortho}} \approx 1 \text{ ps}$

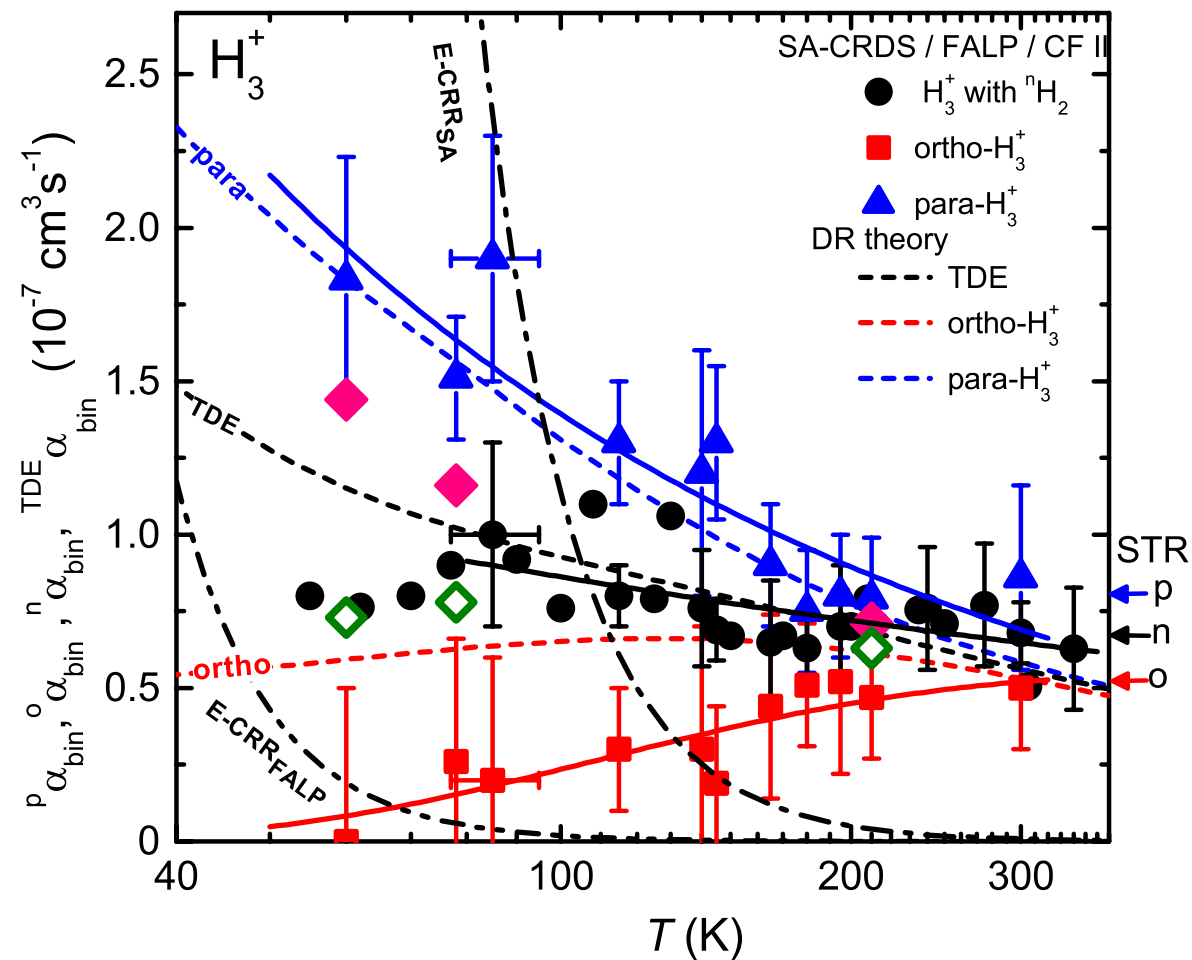
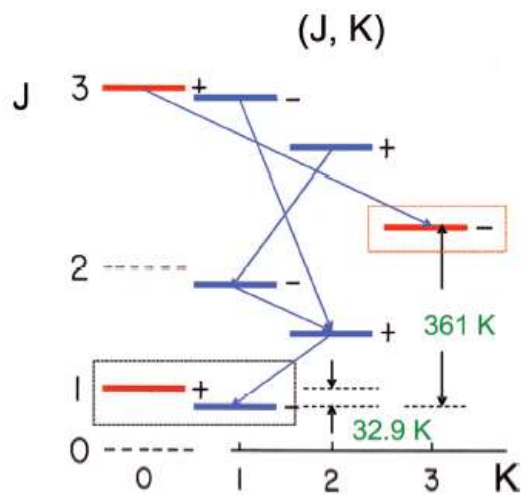
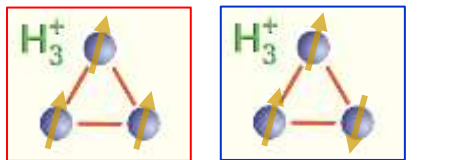


Kokoouline V. and Greene Ch. *Phys. Rev. A* 68 (2003) 012703.
doi: 10.1103/PhysRevA.68.012703

Glośik J. et al. *Phys. Rev. A* 79 (2009) 052707.
doi: 10.1103/PhysRevA.79.052707



Binary DR of para/ortho H_3^+



Hejduk M. *J. Chem. Phys.* 143 (2015) 044303
 doi: 10.1063/1.4927094

