Negative ions





UFP ZS 2024 8C 06.01.2025

Date: 23 Apr 2010 Satellite: Hubble Space Telescope Depicts: Detail of the Carina Nebula



The chemical physics of negative ions

A) Negative ion formation mechanisms: (molecule XYZ)



dissociative attachment

B) Negative ion destruction mechanisms

$$X^- + Y^+ o X^* + Y^*$$
 mutual neutralization

$$X^- + h \upsilon \rightarrow X + e^-$$
 photodetactment

$$X^- + Y o X + Y + e^-$$
 collisional detachment

The negative hydrogen ion H⁻

- one of the most important negative ions in the universe!
- It exists, electron affinity (binding energy of the extra electron) = 0.75 eV

- 🔵

- why does it exist? first electron in H only partially shields the nuclear charge
- QM calculations confirm this
- responsible for most of the continuum opacity of the photosphere

H⁻ in the photosphere

- photosphere what you see when you look at the sun
- about 400 km thick, cool ~ 4400K 5800K, mostly H
- remarkably opaque at infrared and shorter wavelengths
- most H in ground state and thus does not contribute much to absorption
- need 13.6 eV (121.6 nm) to get H in first excited state
- 1939- about one in 10⁷ H's are H⁻, and need only 0.75 eV to remove extra electron → 1653 nm (Saha relation)
- so H⁻ can account for absorption down to very long wavelengths
- negative H makes photosphere as opaque as a dense object, therefore it radiates like a blackbody

negative ions in the earth's ionosphere

- negative ions (O₂⁻⁾ are generally present in the lower ionosphere (D region) 60 – 90 km
- they may play a role in the creation and destruction of the ozone layer observed at 76 km in the polar region

Data from rocket borne instruments



• Neutral density exceeds the ion density below about 500 km.





Ionosphere composition



Dayside ionosphere composition at solar minimum.

Major F-region ions is O^+ , followed by H⁺ at the top and NO⁺ and O₂⁺ at the bottom. Note that neutral gas concentration at 300 km is around 10⁸ cm⁻³, so ion concentrations are 2 orders of magnitude smaller. Negative ions are found only in the lower ionosphere (D region). The net charge of the ionosphere is zero.



- The extent of the ionosphere
 - There are ions and electrons at all altitudes in the atmosphere.
 - Below about 60km the charged particles do not play an important part in determining the chemical or physical properties of the atmosphere.
 - Identification of ionospheric layers is related to inflection points in the vertical density profile.



Negative ions in ionosphere

Negative ions are a feature of the D-region ionosphere, where they hold a substantial portion of the negative charge. Negative ion chemistry is initiated by electron attachment to molecular oxygen

$$O_2 + e + M \rightarrow O_2^- + M \tag{1}$$

after which subsequent reactions form other ions, including complex clusters (see, e.g. Hargreaves, 1992, 231–233). Based on laboratory work and in-situ measurements, the main negative ions are expected to be CO_3^- and NO_3^- , and their hydrates. The main reaction path leading from the initial O_2^- to these "terminal" ions involves neutral ozone, carbon dioxide, and nitric oxide, and the formation of intermediate ion O_3^- (Reid, 1987). Negative ions are present at altitudes below 80 km, where the atmospheric density is high enough so that the 3-body reaction of Eq. (1) is efficient. The balance with electrons is then determined by electron detachment reactions, such as

$$O_2^- + O \rightarrow O_3 + e.$$

Most of the balancing reactions depend on the solar light, such that at night the electrons nearly disappear from altitudes below 80 km and negative charge is held largely by the ions. During sunset, there is a transition of negative charge from electrons to negative ions, and a reverse transition occurs during sunrise. Any realistic modelling of the D-region ionosphere requires consideration of negative ion chemistry.



#	Reaction			Rate coefficient
R1	O ⁻⁺ M	\rightarrow	O+M+e	< 10 ⁻¹⁸
R2	$O_2^-+O_2(^1\Delta_g)$	\rightarrow	2O ₂ +e	2.0×10^{-16}
R3	02+0	\rightarrow	O ₃ +e	1.5×10^{-16}
R4	$O_2^2 + hv(< 2695 \text{ nm})$	\rightarrow	O ₂ +e	3.8×10^{-1}
R5	$NO_{3}^{-}+hv(<318 \text{ nm})$	\rightarrow	NO3+e	5.2×10 ⁻²
R6	$CO_3^2 + hv(<428 \text{ nm})$	\rightarrow	CO ₃ +e	2.2×10^{-12}



Fig. 1. Block diagram for the negative ion scheme of the SIC model. The various ions in the blocks are either the reactants or the final products of the reactions sketched by the connecting arrows. The arrows are labelled with the neutral constituents taking part in the reactions. Note that photodetachment and photodissociation reactions are not shown.

Annales Geophysicae, 24, 187-202, 2006 SRef-ID: 1432-0576/ag/2006-24-187 © European Geosciences Union 2006

(2)





Sunset transition of negative charge in the D-region ionosphere during high-ionization conditions

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- The D Region
 - The most complex and least understood layer in the ionosphere.
 - The primary source of ionization in the D region is ionization by solar X-rays which ionize both N_2 and O_2
 - Lyman- α ionization of the NO molecule.
 - Precipitating magnetospheric electrons may also be important.
 - Initial positive ions are N_2^+ , O_2^+ and NO^+

 $N_2^+ + O_2 \rightarrow O_2^+ + N_2$

- The primary positive ions are O_2^+ and NO⁺
- The most common negative ion is NO₃⁻ The first step in making a negative ion is $e + O_2 + M \rightarrow O_2^- + M$

Negative ions

Formation of negative ions - attachment

Interaction of electrons with atoms and molecules

Photoelectron spectrometer

Photoelectron spectrum



Date: 23 Apr 2010 Satellite: Hubble Space Telescope Depicts: Detail of the Carina Nebula



De Broglie wave length

$$\lambda = \frac{h}{p} = \frac{h}{mv} \sqrt{1 - \frac{v^2}{c^2}}$$

$$\lambda_e(4K) \sim 540 A \sim 54 x 10^{-9} m$$

$$\lambda_e(1eV) \sim 11.6A \sim 1.16x 10^{-9} m$$

$$\lambda_e(mK) \sim 3x10^4 A \sim 3x10^{-6} m$$

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Kvantovka na každý deň

 \mathbf{r}_1

PRL 98, 133201 (2007)

$\frac{H^+ + e^- + e^-}{H^+ + e^-}$





 \mathbf{r}_{2}

We consider the simplest TBR in the case of hydrogen formation, in which two free electrons interact with a proton. To investigate the three-body interaction dynamics, we numerically solve the six-dimensional (6D) timedependent Schrödinger equation, which has the following form (atomic units are used throughout):

$$i\frac{\partial}{\partial t}\Phi(\mathbf{r}_{1},\mathbf{r}_{2},t) = \left[-\frac{1}{2}(\Delta_{\mathbf{r}_{1}}+\Delta_{\mathbf{r}_{2}})-\frac{1}{r_{1}}-\frac{1}{r_{2}} + \frac{1}{|\mathbf{r}_{1}-\mathbf{r}_{2}|}\right]\Phi(\mathbf{r}_{1},\mathbf{r}_{2},t), \quad (1)$$

where \mathbf{r}_1 and \mathbf{r}_2 are the position vectors of each electron, with respect to the proton. We obtain a more tractable with respect to the proton. We obtain a more tractable solution by using the close-coupling recipe [12]: expanding the 6D wave function $\Phi(\mathbf{r}_1, \mathbf{r}_2|t)$ in terms of bipolar spherical harmonics $Y_{l_1 l_2}^{LS}(\Omega_1, \Omega_2)$, $\Phi(\mathbf{r}_1, \mathbf{r}_2|t) =$ $\sum_{LS} \sum_{l_1 l_2} [\Psi_{l_1 l_2}^{(LS)}(r_1, r_2|t)/r_1 r_2] Y_{l_1 l_2}^{LS}(\Omega_1, \Omega_2)$, for a specific symmetry (*LS*). We can also expand the Coulomb repulsion term $1/|\mathbf{r}_1 - \mathbf{r}_2|$ in terms of spherical harmonics. Substituting these expansions into the above Schrödinger Eq. (1) and integrating over the angles Ω_1 and Ω_2 yields a set of coupled partial differential equations with only two radial variables r_1 and r_2 left:

$$\begin{split} i\frac{\partial}{\partial t}\Psi_{j}(r_{1},r_{2}|t) &= [\hat{T}_{1}+\hat{T}_{2}+\hat{V}_{c}]\Psi_{j}(r_{1},r_{2}|t) \\ &+ \sum_{k}\hat{V}_{j,k}^{I}(r_{1},r_{2}|t)\Psi_{k}(r_{1},r_{2}|t), \quad (2) \end{split}$$

where the partial-wave index j runs from 1 to the total number N of partial waves used for expansion. In Eq. (2),

Kvantovka na každý deň

 $i\frac{\partial}{\partial t}\Psi_{j}(r_{1}, r_{2}|t) = [\hat{T}_{1} + \hat{T}_{2} + \hat{V}_{c}]\Psi_{j}(r_{1}, r_{2}|t) + \sum_{k} \hat{V}_{j,k}^{I}(r_{1}, r_{2}|t)\Psi_{k}(r_{1}, r_{2}|t), \quad (2)$

$H^+ + e^- + e^- \rightarrow H + e^-$

$$P_{nl}(E_2) = 2\sum_{LS} \sum_{l_2} \left| \int dr_1 \int dr_2 \phi_{nl}^*(r_1) \phi_{k_2 l_2}^*(r_2) \Psi_{ll_2}^{(LS)}(r_1, r_2, t = t_f) \right|^2,$$

$K_{\rm E} = 0.1 \, {\rm eV}$



FIG. 1 (color online). Snapshots of electron probability distribution on the plane spanned by the radial coordinates r_1 and r_2 for different times: (a) t = 0.0 fs, (b) t = 60 fs, (c) t = 100 fs, (d) t = 150 fs, (e) t = 194 fs, and (f) (in log scale) t = 260 fs.

 r_1 - r_2



Thus, for the case of $K_E = 0.1$ eV considered in Figs. 1 and 2, the total system energy is about $E_{tot} \sim 0.12$ eV instead of $2K_E$. Hence, when one electron recombines to the 10*d* state ($|E_{10d}| \approx 0.136$ eV) of the H atom, the outgoing electron takes an initial total energy of 0.12 eV plus $|E_{10d}|$, thereby $P_{10d}(E_2)$ peaks at $E_2 \sim 0.256$ eV, as shown by the (red) solid line of Fig. 2. Similar energy conservation is also well satisfied for the recombination to the 6*p* state, as is illustrated by the (blue) dash-dotted line in Fig. 2. Our quantum calculations unambiguously reveal the essential feature of a TBR process.

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 $\mathbf{H}^{+} + \mathbf{e}^{-} + \mathbf{e}^{-} \rightarrow \mathbf{H} + \mathbf{e}^{-}$

K_E=0.1 eV







FIG. 3 (color online). The recombination probability P_n as a function of the energy level *n*, for different electron kinetic energies K_E marked in each panel.



FIG. 4 (color online). The recombination probability $P_{n=25,l}$ as a function of the angular-momentum quantum number l, for different electron kinetic energies K_E marked in each panel.



THE ASTROPHYSICAL JOURNAL, 652: L141–L144, 2006 December 1 © 2006. The American Astronomical Society. All rights reserved. Printed in U.S.A.

2006

LABORATORY AND ASTRONOMICAL IDENTIFICATION OF THE NEGATIVE MOLECULAR ION C6H-

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ABSTRACT

The negative molecular ion C_6H^- has been detected in the radio band in the laboratory and has been identified in the molecular envelope of IRC +10216 and in the dense molecular cloud TMC-1. The spectroscopic constants derived from laboratory measurements of 17 rotational lines between 8 and 187 GHz are identical to those derived from the astronomical data, establishing unambiguously that C_6H^- is the carrier of the series of lines with rotational constant 1377 MHz first observed by K. Kawaguchi et al. in IRC +10216. The column density of C_6H^- toward both sources is 1%–5% that of neutral C_6H . These surprisingly high abundances for a negative ion imply that if other molecular anions are similarly abundant with respect to their neutral counterparts, they may be detectable both in the laboratory at high resolution and in interstellar molecular clouds.

Note added in proof.—A third member in the series, C_4H^- , has now been detected in the laboratory at centimeter and millimeter wavelengths. A full account of this work will be presented elsewhere.

Interstellar Molecules

2 atoms	3 atoms	4 atoms	5 atoms	6 atoms	7 atoms	8 atoms	9 atoms	10 atoms	11 atoms	12 atoms	13 atoms
H ₂	C ₃ *	c-C₃H	C ₅ *	C₅H	C ₆ H	CH ₃ C ₃ N	CH₃C₄H	CH ₃ C₅N	HC ₉ N	C ₆ H ₆	HC ₁₁ N
AIF	C₂H	I-C₃H	C₄H	I-H ₂ C ₄	CH ₂ CHCN	HC(O)OCH ₃	CH ₃ CH ₂ CN	(CH ₃) ₂ CO	CH₃C ₆ H	C ₂ H ₅ OCH ₃	
AICI	C ₂ O	C₃N	C₄Si	C ₂ H ₄ *	CH ₃ C ₂ H	CH₃COOH	(CH ₃) ₂ O	(CH2OH) ₂	C₂H₅OCHO	n-C ₃ H ₇ CN	
C2**	C ₂ S	C₃O	I-C ₃ H ₂	CH₃CN	HC₅N	C ₇ H	CH ₃ CH ₂ OH	CH₃CH₂CHO			-
СН	CH2	C₃S	c-C ₃ H ₂	CH₃NC	CH₃CHO	H ₂ C ₆	HC ₇ N		-		
CH⁺	HCN	$C_{2}H_{2}^{*}$	H2CCN	CH₃OH	CH ₃ NH ₂	CH₂OHCHO	C₅H				
CN	НСО	NH₃	CH ₄ *	CH₃SH	c-C₂H₄O	I-HC ₆ H*	CH ₃ C(O)NH ₂				
СО	HCO⁺	HCCN	HC₃N	HC ₃ NH ⁺	H₂CCHOH	CH₂CHCHO	C₅H⁻				
CO ⁺	HCS ⁺	HCNH ⁺	HC₂NC	HC ₂ CHO	C₀H⁻	CH₂CCHCN	C ₃ H ₆				
СР	HOC ⁺	HNCO	нсоон	NH ₂ CHO		H ₂ NCH ₂ CN		-			
SiC	H₂O	HNCS	H ₂ CNH	C₅N			-				
HCI	H₂S	HOCO ⁺	H ₂ C ₂ O	I-HC₄H*		Cations					
KCI	HNC	H ₂ CO	H ₂ NCN	I-HC₄N							
NH	HNO	H ₂ CN	HNC ₃	c-H ₂ C ₃ O		Anions					
NO	MgCN	H₂CS	SiH ₄ *	H₂CCNH							
NS	MgNC	H₃O ⁺	H ₂ COH ⁺	C₅N [−]	8 -	1 1	1 1 1				
NaCl	N_2H^+	c-SiC₃	C₄H⁻		2954	3220	3715 Q	4489 4685 CCN 4848 Q 4872	5107 5145 N		
ОН	N ₂ O	CH ₃ *	HC(O)CN		- U105,	- U106.	- C2H5C	- U106. - H ¹³ C	- U106.		
PN	NaCN	C₃N⁻			4 –						ei et al.,
SO	OCS	PH3						A		ApJ 6	00, 234
SO⁺	SO ₂	HCNO			2	, ind	NL A	ma h] (20	004)
SiN	c-SiC ₂	HOCN			0 perfusion	when non when when	1 Malana Marily	W. MM W V		WA .	/
SiO	CO ₂ *	HSCN							BIMA (smooth	ied)	
SiS	NH2				2	i ı	1 1 1	1 1			
CS	$H_3^{\dagger *}$				106.	3 106.325 106.35 1	06.375 106.4 106.42	5 106.45 106.475 10	6.5 106.525	JAR .	
HF	H_2D^+)					Frequenc	y (GHZ)		The to	
SH*	HD ₂ ⁺								11 - N		JR (
HD	SiCN										
FeO	AINC									100%	AL
0 ₂	SINC									70-01	1
CF	HCP	https://cdms.astro.uni-koeln.de/classic/molecules									
	CUP									1 2 2 2 1 2 1 2 1 2 1 2 1 2 1 2 1 2 1 2	
PU											**

AIO

2 atoms	3 atoms	4 atoms	5 atoms	6 atoms	7 atoms	8 atoms	9 atoms	10 atoms	11 atoms	12 atoms	>12 atoms
CH+	C ₃ *	c-C ₃ H	C ₅ *	C ₅ H	C ₆ H	CH ₃ C ₃ N	CH ₃ C ₄ H	CH ₃ C ₅ N	HC ₉ N	c-C ₆ H ₆ *	C ₆₀ *
СН	C ₂ H	I-C ₃ H	C ₄ H	I-H ₂ C ₄	CH ₂ CHCN	HC(O)OCH ₃	CH ₃ CH ₂ CN	(CH ₃) ₂ CO	CH ₃ C ₆ H	n-C ₃ H ₇ CN	C70*
CN	C ₂ O	C ₃ N	C ₄ SI	C ₂ H ₄ *	CH ₃ C ₂ H	СН3СООН	(CH ₃) ₂ O	(CH ₂ OH) ₂	C ₂ H ₅ OCHO	I-C ₃ H ₇ CN	C ₆₀ **
он	C ₂ S	C ₃ O	I-C ₃ H ₂	CH ₃ CN	HC ₅ N	C ₇ H	CH ₃ CH ₂ OH	CH ₃ CH ₂ CHO	CH ₃ OC(O)CH ₃	C ₂ H ₅ OCH ₃	c-C ₆ H ₅ CN
co	CH ₂	C ₃ S	c-C ₃ H ₂	CH ₃ NC	CH ₃ CHO	C ₆ H ₂	HC ₇ N	CH ₃ CHCH ₂ O	CH ₃ C(O)CH ₂ OH	1-c-C ₅ H ₅ CN	HC ₁₁ N
H ₂	HCN	C ₂ H ₂ *	H ₂ CCN	CH ₃ OH	CH ₃ NH ₂	CH ₂ OHCHO	C ₈ H	CH ₃ OCH ₂ OH	c-C ₅ H ₆	2-c-C ₅ H ₅ CN	1-C ₁₀ H ₇ CN
SiO	HCO	NH ₃	CH4*	CH ₃ SH	c-C ₂ H ₄ O	I-HC ₆ H*	CH ₃ C(O)NH ₂	c-C ₆ H ₄	HOCH ₂ CH ₂ NH ₂	CH ₃ C ₇ N (?)	2-C ₁₀ H ₇ CN
CS	HCO*	HCCN	HC ₃ N	HC ₃ NH ⁺	H ₂ CCHOH	CH ₂ CHCHO	C ₈ H ⁻	H ₂ CCCHC ₃ N	H ₂ CCCHC ₄ H	n-C ₃ H ₇ OH	c-C ₉ H ₈
SO	HCS*	HCNH*	HCCNC	HCCCHO	C ₆ H ⁻	CH ₂ CCHCN	C ₃ H ₆	C ₂ H ₅ NCO	C ₁₀ H ⁻ (2023)	i-C ₃ H ₇ OH	1-c-C5H5CCH
SiS	HOC+	HNCO	нсоон	NH ₂ CHO	CH ₃ NCO	H ₂ NCH ₂ CN	CH ₃ CH ₂ SH	C ₂ H ₅ NH ₂ (?)	H ₂ C(CH) ₃ CN (2023)	(CH ₃) ₂ C=CH ₂ (2023)	2-c-C ₅ H ₅ CCH
NS	H ₂ O	HNCS	H ₂ CNH	C ₅ N	HC ₅ O	CH ₃ CHNH	CH ₃ NHCHO	HC ₇ NH ⁺			c-C ₅ H ₄ CCH ₂
C2**	H ₂ S	HOCO ⁺	H ₂ C ₂ O	I-HC ₄ H*	HOCH ₂ CN	CH ₃ SiH ₃	HC ₇ O	E-CH ₃ CHCHCN			2-C ₉ H ₇ CN
NO	HNC	H ₂ CO	H ₂ NCN	I-HC ₄ N	HCCCHNH	H ₂ NC(0)NH ₂	HCCCHCHCN	Z-CH ₃ CHCHCN			C ₆ H ₅ CCH (2023)
HCI	HNO	H ₂ CN	HNC ₃	c-H ₂ C ₃ O	HC ₄ NC	HCCCH ₂ CN	H ₂ CCHC ₃ N	CH ₃ C(CN)CH ₂			CH ₃ OCH ₂ CH ₂ OH (2024)
NaCl	MgCN	H ₂ CS	SiH4*	H ₂ CCNH	c-C3HCCH	HC ₅ NH*	H ₂ CCCHCCH	CH2CHCH2CN			1-C ₁₂ H ₇ CN (2024)
KCI	MgNC	H ₃ O ⁺	H ₂ COH*	C ₅ N ⁻	I-H ₂ C ₅	CH ₂ CHCCH	HOCHCHCHO (2024)	HOCH ₂ C(O)NH ₂ (2023)			5-C ₁₂ H ₇ CN (2024)
AICI	N_2H^*	c-SIC ₃	C ₄ H ⁻	HNCHCN	MgC ₅ N	MgC ₆ H	HC ₇ N ⁺ (2024)	CH ₃ CH ₂ CCH (2024)			1-C ₁₆ H ₉ CN (2024)
AIF	N ₂ O	CH3*	HC(O)CN	SIH ₃ CN	CH ₂ C ₃ N	C ₂ H ₃ NH ₂	CH ₂ (CCH) ₂ (2024				
PN	NaCN	C ₃ N ⁻	HNCNH	C ₅ S	NC ₄ NH ⁺ (2023)	(CHOH) ₂					
SIC	OCS	PH ₃	CH ₃ O	MgC ₄ H	MgC ₅ N ⁺ (2023)	HC ₂ (H)C ₄					
CP	SO ₂	HCNO	NH4 ⁺	CH ₃ CO ⁺	HC ₅ N ⁺ (2024) C ₇ N ⁻ (2023)					
SIN	CO ₂ *	HSCN	NCCNH*	H ₂ C ₃ S	CH ₂ (CN) ₂ (2024	MgC ₆ H ⁺ (2023)					
SO ⁺	NH ₂	H ₂ O ₂	CH ₃ CI	HCCCHS		Z-(CH) ₂ (CN) ₂ (2024		"m	ore th	an 320	molec
0*	H3*(*)	C ₃ H ⁺	MgC ₃ N	C ₅ O							
F	SICN	HMgNC	NH ₂ OH	C ₅ H*							
iH?	AINC	нссо	HC ₃ O*	HCCNCH*							
eO?	SINC	CNCN	HC ₃ S*	c-C ₃ C ₂ H							
02	HCP	HONO	H ₂ C ₂ S	HC ₄ S				f Novon	abor 2	074 27	7 molo
F+	CCP	MgC ₂ H	C ₄ S	HMgC ₃ N (2023)			AS 0	1 NOVEL		024, 32	
*0	AIOH	HCCS	HC(O)SH	MgC ₄ H ⁺ (2023)							
NO	H ₂ O*	HNCN	HC(S)CN	H ₂ C ₃ H ⁺ (2023)							
DH+	H ₂ CI ⁺	H ₂ NC	нсссо	H ₂ C ₃ N (2023)							
N.	KCN	HCCS*	NaCCCN (2023)	(HO) ₂ CO (2023)							
SH*	FeCN	CH3 ⁺ (2023)	MgC ₃ N ⁺ (2023)	H ₂ CNCN (2024)							
SH	HO ₂	HCNS (2024)	HC ₃ N ⁺ (2024)	NCHCCS (2024)							
HCI*	TIO ₂	HOCS ⁺ (2024)	HC ₃ S (2024)								
10	C ₂ N	HNSO (2024)	NC ₃ S (2024)								
vH+	Si ₂ C										
2	HS ₂										
0*	HCS										
S*	HSC										
leH*	NCO										
PO+	CaNC										
SIP ?	NCS										
(2023)	mguz										
MgS (2024)	HSO										
(2024)	(2023)										

.

Extragalactic Molecules (as of 03/2024)

2 atoms	3 atoms	4 atoms	5 atoms	6 atoms	7 atoms	8 atoms	>8 atoms
он	H ₂ O	H ₂ CO	c-C ₃ H ₂	CH ₃ OH	CH3CCH	HC ₆ H	c-C ₆ H ₆ *
со	HCN	NH ₃	HC ₃ N	CH ₃ CN	CH ₃ NH ₂	HC(O)OCH ₃	C ₆₀ *(?)
H ₂ *	HCO*	HNCO	CH ₂ NH	HC ₄ H*	CH ₃ CHO		(CH ₃) ₂ O
СН	C ₂ H	C ₂ H ₂ *	NH ₂ CN	HC(O)NH ₂	C ₂ H ₃ CN		
CS	HNC	H ₂ CS?	I-C ₃ H ₂	CH ₃ SH			
CH ⁺	N_2H^+	HOCO*	H ₂ CCN				
CN	ocs	c-C ₃ H	H ₂ CCO				
SO	нсо	H ₃ O ⁺	C ₄ H				
SiO	H ₂ S	I-C ₃ H	НСООН				
CO ⁺	SO ₂	C ₃ N	C ₂ H ₃ ⁺ (2024)				
NO	HOC*	I-C ₃ H ⁺					
NS	C ₂ S	H ₂ CN					
NH	H_2O^+	H ₂ NC					
OH ⁺	HCS ⁺						
HF	H ₂ Cl ⁺						
SO*	NH ₂						
ArH ⁺							
CF ⁺							
SH*							
O ₂ (?)							
PN							

Very low collision energies

 $\operatorname{Kr}({}^{1}S_{0}) + \hbar\omega \rightarrow \operatorname{Kr}^{+}({}^{2}P_{1/2}) + e^{-}(\varepsilon)$

 $e^- + SF_6 \rightarrow SF_6^{-*}$

 $e^+ + CCl_4 \rightarrow CCl_4^{-*} \rightarrow CCl_3 + Cl^-$



Figure 1. Schematic diagram of the vuv photoionization apparatus used for attachment studies (Chutjian and Alajajian 1985a, b).



TOPICAL REVIEW

J. Phys. B: At. Mol. Opt. Phys. 28 (1995) 1645-1672. Printed in the UK

Electron-molecule collisions at very low electron energies

F B Dunning

Department of Physics and the Rice Quantum Institute, Rice University, PO Box 1892, Houston, TX 77251, USA

1995



Interaction of electrons with atoms and molecules Formation of negative ions <u>attachment</u>

Existence of negative ions

It is in case of H- second electron in the field of proton and first electron



+e
$$-\alpha e$$

 \downarrow \Box
 $V_{\text{within}} = -(1-\alpha)e/r_1$

$$\beta < 1$$

 $V_{out} = \beta(1-\alpha)e/r_1$

Total potential energy of electron of the charge -e

$$U(r_1)_{total} = -(1-\beta)(1-\alpha)e^2/r_1$$

Calculation of H⁻ potential

The calculation of the potential energy of an electron in the field of a hydrogen atom may be carried out to a first approximation as follows. If we neglect polarization effects (which are actually decisive in determining the stability of H⁻) the probability of finding the atomic electron at a distance between r and r + drfrom the nucleus is $4\pi r^2 \psi^2 dr$, where $\psi = (\pi a_0^3)^{-1/2} \exp(-r/a_0)$ is the wave function of the ground state of hydrogen. The potential energy due to the atomic electron at a point distant r_1 from the nucleus is

$$\frac{4\pi e^2}{r_1} \int_0^{r_1} \psi^2 r^2 \,\mathrm{d}r + 4\pi e^2 \int_{r_1}^\infty \frac{\psi^2 r^2}{r} \,\mathrm{d}r,$$

the first term arising from the charge within r_1 , the second from that without. Carrying out the elementary integrations involved gives

$$e^{2}\left(\frac{\mathbf{I}}{r_{1}}-\exp(-2r_{1}/a_{0})\left(\frac{\mathbf{I}}{r_{1}}+\frac{\mathbf{I}}{a_{0}}\right)\right)$$

Adding the potential energy $-e^2/r_1$ due to the nucleus we find for the total potential energy

$$-e^2 \exp(-2r_1/a_0) \left(\frac{\mathbf{I}}{r_1} + \frac{\mathbf{I}}{a_0}\right) \cdot$$
 (1.1)



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- The D Region
 - The most complex and least understood layer in the ionosphere.
 - The primary source of ionization in the D region is ionization by solar X-rays which ionize both N_2 and O_2
 - Lyman- α ionization of the NO molecule.
 - Precipitating magnetospheric electrons may also be important.
 - Initial positive ions are N_2^+ , O_2^+ and NO^+

 $N_2^+ + O_2 \rightarrow O_2^+ + N_2$

- The primary positive ions are O_2^+ and NO⁺
- The most common negative ion is NO₃⁻
 The first step in making a negative ion is

 $e + O_2 + M \rightarrow O_2^- + M$

Attachment (to molecule) and detachment



Fig. 1.1. Born-Oppenheimer potential energy curves illustrating electron affinity (EA) and vertical detachment energy (VDE).





Fig. 9.1. Potential-energy curves illustrating three possible ways in which negative ions may be formed from a molecule XY by electron capture.

Photodetachment $O_2^- + hv \rightarrow O_2 + e$

O (2P)+O(1D)

O(3P)+O(3P)

0(2P)+O(3P)

Fig. 11.8 shows a typical arrangement used in the study of photodetachment from NO⁻ and O₇ which is very similar to that used for He⁻.



and energy distribution of electrons resulting from photodetachment by laser light. From Siegel et al. (1972). Do, D1, D2, vertical and horizontal deflectors; L_0, L_1, L_2 einzel lenses; Q_1, Q_2, Q_3, Q_4 , twelve-element symmetrical quadrupole lenses.





Fig. 11.32. Energy spectrum of the photoelectrons arising through photodetachment from O_2^- by argon ion laser light, as observed by Celotta et al. (1972).

Vibrational predissociation

Figure 1.10 illustrates vibrational predissociation. The coordinates Q_1 and Q_2 represent different molecular motions in a polyatomic molecule. The anion M⁻ formed on electron attachment has sufficient energy to dissociate but the energy is initially in vibrational modes that do not correspond to the reaction coordinate Q_2 .



Fig. 1.10. Schematic potential energy curves illustrating dissociative electron attachment via vibrational predissociation. Q_1 and Q_2 represent different motions in a polyatomic ion.

Hydrogen molecule



The discovery of large concentrations of H– ions (30%) in low-temperature hydrogen plasmas [1] has provided the possibility of producing intense beams of H– ion sources for neutral beam heating [2, 3] of thermonuclear fusion devices [4, 5]. Modelling of such plasmas requires a detailed knowledge of the mechanism of dissociative electron attachment (DEA) of electrons to molecular hydrogen. Schematically the process may be written as a two-step process:

$e^{-} + H_2(v, J) \rightarrow H_2^{-} \rightarrow H^{-} + H$

where v and J denote the vibrational and rotational quantum numbers of the target hydrogen molecule. The incident electron is trapped by the neutral molecule in a resonant state, this trapping increasing the electron residence time sufficiently to allow the relatively slow moving nuclei to dissociate. In the present case the $H_2^-(1\sigma_g^2 1\sigma_u^2 + 2\Sigma_u^+)$ resonance formed at a range of electron energies of around 4 eV is characterized by a very short lifetime against autodetachment when formed from the v = 0 level. Hence the cross section for product H– is low.

Hydrogen molecule EA calculated by J. Horacek



Hydrogen molecule







Fig. 9.20. Variation with electron energy of negative-ion production by electron impact in H_2 , observed by Schulz (1959). ---- additional production observed when the reagent grade hydrogen was introduced without a liquid air trap.

Vibr. excitation of N₂ fine structure







of a theory such as that outlined above. Haas suggested that we must regard the collisions as taking place in two stages—the incident electron is first captured to form a negative ion N_2^- that is energetically unstable but has a lifetime greater than a vibrational period. It eventually breaks up, becoming a neutral molecule that may be in an excited vibrational state—in other words, the process is regarded as a resonance one of the same type as that found in elastic scattering of electrons by helium and other atoms and molecules (see Chap. 9).



FIGURE 1. Potential energy curves for N1 and N1+.+



FIG. 11.30. Momentum-transfer cross-section for electrons in N_2 . —— derived by Engelhardt, Phelps, and Risk from analysis of swarm data. — · — · — · — derived by Pack and Phelps from analysis of their drift velocity observations. — — derived from drift velocity observations of Crompton and Sutton. · · · · total cross-section measured by Ramsauer method.



FIG. 11.31. Cross-sections for rotational and vibrational excitation of nitrogen. Q_4^6 is the cross-section for the rotational excitation $J = 4 \rightarrow J = 6$. $\Sigma_v Q_v$ is the sum of the cross-sections for vibrational excitation consistent with the swarm data.

Rotational excitation N₂





Figure 2-2-1. Vibrational-rotational levels (quantum numbers v and J) of a few diatomic molecules. The (v = 1, J = 0) level of H₂ lies 0.54 eV above the ground state (v = 0, J = 0). Rotational level spacings for H₂ are uniquely large, about 15J meV, where J is the quantum number for the upper level. For the ortho species of H₂ $(o-H_2)$, the nuclear spins are parallel; for the para version $(p-H_2)$, the nuclear spins are antiparallel. [From Shimamura (1984).]

Cross section of electron attachment – idea of experiment


Electron attachment to HBr



EA to CCl_2F_2 absolute cross section



Fig. 1. Absolute electron attachment cross section vs. electron energy for the reaction $CCl_2F_2 + e \rightarrow Cl^-$. The present data (full line with full squares) were derived by using the calibration method outlined in the text [Eqs. (4) and (5)] involving integration of the measured anion signal. The earlier data (dashed line) are from [17] and have been obtained by [17] using the simple procedure employing Eq. (7). Also shown as full line designated ϵ^{-1} the predicted energy dependence for *s*-wave scattering.

Cross sections comparison





Calculated DA cross section in HI (*full curve*). *Diamonds* represent the experimental data of Klar et al. [24]



Attachment HI 20meV

Electron Attachment Spectroscopy (EAS)



Electron Attachment Spectroscopy – electron energy distribution



Three-body attachment $F_6 + e \rightarrow SF_6^-$ SF₆ + e + M $\rightarrow SF_6^-$ + M



Fig. 2.2. Ion yield curve for SF_6^{-*} formation and electron beam intensity monitored at collector A.

SF_ + F + F⁻ SF_ + F₂ SF_ + F₂⁻

– SFς + F⊺

EA SF₆





EA SF₆ temperature dependence



Figure 6-8-9. (a) Schematic potential energy curves for SF_6 and SF_6^- . (b) Electron attachment to form SF_5^- from SF_6 versus electron energy, for various gas temperatures. [From Chen and Chantry (1979).]

Attachment NO to particular channel



Fig. 5. Schematic potential energy curves for the ground states of NO and NO⁻ and two exited states of NO^{*-} after Ref. [8]. Also shown is the relative DA cross-section curve for the reaction NO+e \rightarrow O⁻(²P)+N^{*}(²D) obtained by applying the reflection principle.



Fig. 1. Schematic diagram of the apparatus. Electrons are emitted from a hot filament, formed into a beam, pass the hemispherical energy selector at a constant energy of ~ 2 eV and are focused and brought to the final collision energy before they interact with the neutral beam and are collected at a Faraday cup.

Dissociative electron attachment to NO close to threshold may in principle proceed via the following channels:

$$NO + e \rightarrow O^{-}(^{2}P) + N(^{4}S),$$
 (1a)

$$NO + e \rightarrow O^{-}(^{2}P) + N^{*}(^{2}D),$$
 (1b)

$$NO + e \rightarrow O^{-}(^{2}P) + N^{*}(^{2}P).$$
 (1c)

The threshold energies of 5.074, 7.457 and 8.650 eV, respectively, can be derived using the bond dissociation energy D(N-O) = 6.535 eV [10], the electron affinity EA(O) = 1.461 eV [11] and the excitation energies $E(N^{*}(^{2}D)) = 2.383$ eV and $E(^{*}N(^{2}P)) =$



Fig. 4. Formation of O⁻ from NO by electron impact. The Cl^-/CCl_4 cross-section curve is used to calibrate the electron energy scale and to determine the electron energy resolution (120 meV). The arrow points to the presently determined threshold energy of the reaction $NO + e \rightarrow O^-(^2P) + N^*(^2P)$.

EA clusters and surface







Energy levels in clusters



Fig. 3. Upper panel: approximate potential energy diagrams for NO and NO⁻ adapted from Ref. [5]. The $\dot{v} = 0$ level is located at 26 meV below the v = 0 level of neutral NO. Lower panel: potential energy curves as above but shifted for the anion due to the polarization interaction of the negative charge with surrounding NO molecules. The NO⁻ curve is positioned (arbitrarily) so that the $\dot{v} = 4$ level is accessible by 40 meV electrons, in accordance with the experiment.



Fig. 1. NO⁻ signal observed from electron attachment to an NO cluster beam as a function of the electron energy. For comparison the energy calibration signal Cl⁻ from CCl₄ is also shown. Insets: count rate versus electron energy on a log-log scale. Dotted lines indicate the 1/E dependence predicted by s-wave scattering theory.

Negative ions, formation of negative clusters



Fig. 1. Schematic diagram of the instrument. Electrons are emitted from a hot filament, and focused into a beam. They pass the hemispherical energy selector at a constant energy of about 4 eV and are focused and brought to the final collision energy before they interact with the neutral beam and are collected at a Faraday cup. Also shown is a close-up of the monochromator in a three dimensional view.



Fig. 2. Mass spectrum for $(N_2O)_nO^-$ cluster anions produced by electron attachment (using 2 eV electrons) to a N₂O cluster beam formed with the stagnation gas temperature at room temperature and the stagnation gas pressure of 3 bar.



Fig. 6. Energy dependence of the $(N_2O)_{\mu}O^-$ yield for various cluster sizes (all curves are normalized to the same maximum value) for an extended range of electron energy (as compared to results shown in Fig. 4 which only cover the range up to 4 eV).

Rydberg atom attachment apparatus



Figure 6-6-2. Dunning's Rydberg atom attachment apparatus. [From Dunning (1987).]

X^{*} are produced by laser excitation

 $\mathbf{X}^* + \mathbf{AB} \rightarrow \mathbf{X}^+ + \mathbf{AB}^-$

 X^+ and AB^- are detected by the pair of opposing detectors. Time-of-flight technique can be used to determine the mass of the negative ions The density of Rydberg atoms in the interaction region can be determined by applying sufficiently strong electric field to ionize the atoms and then counting the ions thus formed.

Electron attachment from laser excited Rydberg atoms - Hotop



Figure 6-6-6. Results from Hotop's experiments on electron attachment from laserexcited Rydberg atoms and photodetached threshold electrons. [From Klar et al. (1991).] X^{*} are produced by laser excitation

X* + AB → X+ + AB-

Photodetachment of negative ions

6-10. DETACHMENT PROCESSES

A. Photodetachment of Negative Ions

The detachment of electrons from negative ions by photon impact has been the subject of intensive investigation since the pioneering work of Branscomb and collaborators in the 1950s and 1960s (Branscomb, 1962). Recent reviews of progress in this area have been given by Miller (1981), Drzaic et al. (1984), Mead et al. (1984), and Ervin and Lineberger (1991). By measuring the location of photodetachment thresholds, it is possible to obtain extremely accurate values of electron affinities for atomic and some molecular negative ions, as discussed in Section 6-2. The high photon flux available from laser light sources has led to improvements in accuracy of many orders of magnitude for such measurements.

A typical apparatus used for photodetachment studies is that shown in Fig. 6-10-1. Negative ions are produced in a hot-cathode arc discharge or cold-



Figure 6-10-1. Apparatus used by Lineberger and Woodward (1970) for studying photodetachment with a tunable dye laser as a light source.

Atomic Collisions – Heavy particle projectiles (1993) E.W.McDaniel, J.B.A.Mitchell, M.E.Rudd,

Threshol Studies



Figure 6-10-3. (a) Photodetachment cross sections for Se⁻, measured with a tunable dye laser as a light source, by Lineberger and Woodward (1970) and by Hotop et al. (1973), respectively. (Not all measured data points are shown. (b) Energy levels of Se and Se⁻.

Cross section

Photodetachment $O^- + hv \rightarrow O + e$



Fig. 11.13. Photodetachment cross-sections for O^* . \bullet , \circ , observed by Branscomb *et al.* (1965), (1) derived from analysis of arc emission spectra by Boldt (1959) (see Chapter 8, p. 253), (2) calculated by Cooper and Martin (1962), (3) calculated by Robinson and Geitman (1967).



FIG. 2. (Color online) Measured cross section of O^- as a function of the photon energy. Our data (large full circles) is compared with the relative measurements of Refs. 13 and 25 (squares and small triangles), which were calibrated to hydrogen anion measurements. The dashed line shows the *ab initio* calculation of Ref. 15.

Photodetachment $OH^- + hv \rightarrow OH + e$

Electron affinity of O is 1.46eV



Figure 6-10-6. Photodetachment cross sections for O⁻. [From Lee and Smith (1979).]



Fig. 3 Photodetachment cross section of OH⁻ at a trap temperature of 50 K for varying photon energy. The steps in the cross section are due to the opening of loss channels corresponding to the J = 2, 1 and 0 rotational states of the anion.



na tomto obrázku data od Mulin 2015 byla korigovaná na T_{22PT} +5 K! Teploty pro data od Mulin 2015 jsou přepočtené na srážkové teploty T = T_{22PT} +5 K.

Photodetachment $O^- + hv \rightarrow O + e$

Photodetachment $OH^- + hv \rightarrow OH + e$



Fig. 3 Photodetachment cross section of OH⁻ at a trap temperature of 50 K for varying photon energy. The steps in the cross section are due to the opening of loss channels corresponding to the J = 2, 1 and 0 rotational states of the anion.

Internal state thermometry of cold trapped molecular anions

Cite this: Phys. Chem. Chem. Phys., 2013, 15, 612

Rico Otto,†‡^{ab} Alexander von Zastrow,†§^b Thorsten Best^a and Roland Wester*^a



Fig. 3 Photodetachment cross section of OH^- at a trap temperature of 50 K for varying photon energy. The steps in the cross section are due to the opening of loss channels corresponding to the J = 2, 1 and 0 rotational states of the anion.



transitions	threshold energies ε _ι (cm ⁻¹)	transitions	threshold energies ε _ι (cm ⁻¹)
O ₃ (4)	14450.8	P ₁ (3)	14703.6
O ₃ (3)	14516.4	Q ₃ (2)	14712.5
P ₃ (4)	14569.1	Q ₃ (3)	14718.8
P ₃ (3)	14600.2	Q ₃ (4)	14723.0
P ₃ (2)	14628.6	R ₃ (0)	14740.9
P ₁ (4)	14655.7	R ₃ (1)	14787.2
Q ₃ (1)	14703.5	Q ₁ (4)	14796.1

 $R_3(0), Q_3(1) \text{ and } P_3(2)$

Table 1. The list of transitions and corresponding threshold energies used in present study. Note that only the transitions $R_3(0)$, $Q_3(1)$ and $P_3(2)$ were considered in the study by Otto et al. (2013).



Fig 2. Experimentally determined relative populations of the lowest rotational states of OH- anion (squares) compared to calculated thermal populations at 116 K (circles). Due to the position of the corresponding photodetachment thresholds, the value obtained for J = 3 consists from indistinguishable contributions from states with J = 3, 4 and 5. Sum of the calculated relative thermal populations of these states is denoted by cross.



Fig 1. Photodetachment rate at trap temperature of 115 K. The obtained rotational temperature is $Trot = 111 \pm 6$ K. Red line is the expected photodetachment cross section at 115 K under assumption of thermal population of OH⁻ rotational states. Green line denotes fit of the data by equation (1) under assumption of thermal population of states with temperature as a free parameter and the blue line denotes fit of the obtained data by equation (1) with the populations of rotational states up to J=3 as free parameters.



Fig. 3 Photodetachment cross section of OH⁻ at a trap temperature of 50 K for varying photon energy. The steps in the cross section are due to the opening of loss channels corresponding to the J = 2, 1 and 0 rotational states of the anion.

Cross sections comparison

1000

50



HI Dissociative attachment Photon Energy (eV) 1.5 1.7 1.9 $O^- + hv \rightarrow O + e$ 0 75 100 125 150 175 200 electron energy E [meV] Cross Section (10⁻¹⁸ cm²) ^P1/2 ^P3/ Lee and Smith Branscomb et al. 8000 7500 7000 6500 8500 Wavelength (eV) Figure 6-10-6. Photodetachment cross sections for O⁻. [From Lee and Smith (1979).]

Attachment

Photodetachment x10⁻²

Photodetachment $H^- + hv \rightarrow H(n) + e$

Cross section



Hanspeter Helm - Photoelectron spectrometer - Experimental setup



Schematical view of the fast negative ion beam imaging spectrometer. A fast beam of negative hydrogen ions is formed in a hollow cathode discharge ion source, and shaped by an einzel lens 1. The einzel lens 1 collimates the 3 keV beam for optimal passage through the Wien-Filter. The settings of the Wien-Filter also prohibit the passage of electrons co-propagating in the beam into subsequent vacuum chambers. The einzel lenss 2 and 3 serve as an electrostatic telescope to control the collimation and convergence properties of the beam in the interaction chamber. The H dux corresponds to a current of 100 nA and arrives at the center of the photoelectron spectrometer with a beam waist of 400 m. The 90 degree beam bend in the quadrupole deflector removes neutral hydrogen atoms from the beam. These are produced by collisional detachment with residual gas atoms in the first two vacuum sections where a higher residual pressure prevails. The three vacuum sections are differentially pumped to maintain a residual pressure of $5 \, 10^{-10}$ mbar in the third section during operation. Intense laser pulses of an energy of 56 J, a pulse length of 250 fs and a wavelength of 2.15 m cross the ion beam under 90 degrees and interact with atomic particles at a repetition rate of 1 kHz. We use a standard Ti:Sapphire laser system with a regenerative amplifier and an optical parametric amplifier (OPA). The outgoing pulses are linearly polarized along the ion beam propagation axis and focused at the center of the ion beam. The heart of our setup is an imaging spectrometer that was first introduced by Helm *et al*, see following box.

Hanspeter Helm Photoelectron spectrometer

The principle of our Photoelectron spectrometer is shown on the right. Electrons produced in the interaction volume by laser irradiation of atoms, ejected in a solid angle of 4π are mapped onto a 2D position sensitive detector. The projection is achieved by homogeneous or weakly inhomogeneous electrical fields of 80 V/cm^2 in the inner spectrometer region. The detector consists of a Chevron stack of 2 inch diameter high-quality Multi-Channel Plates (MCPs) and a phosphor screen coated by a transparent, conducting gold layer. The amplified signal of an electron impact is drawn onto the phosphor screen by a potential of 2 keV to enhance in phosphorescence yield. This finite-sized light spots are accumulated and integrated by a 12 bit charge-coupled-device (CCD) camera and the data are taken by a frame grabber. The total electron yield is 10^7 - 10^8 spatially resolved electrons per hour acquisition time. Conventional photoelectron spectroscopy usually requires higher repetition laser systems in order to achieve similar statistics.



Hanspeter Helm 3D



Projection of a continuous 3D electronic wavefunction in momentum space onto a 2D imaging detector. It consists of a mixture of different partial waves and discrete energies. Their energetic spacing corresponds to the photon energy of the laser (EPD Excess Photon Detachment). The modulus of the 3D wavefunction can be reconstructed using numerical inversion routines, since dipole transitions possess an intrinsic cylindrical symmetry.

22 pole trap

THE JOURNAL OF CHEMICAL PHYSICS 130, 061105 (2009)

Absolute photodetachment cross section measurements of the O⁻ and OH⁻ anion

P. Hlavenka, R. Otto, S. Trippel, J. Mikosch,^{a)} M. Weidemüller,^{b)} and R. Wester^{c)}

Absolute total photodetachment cross sections of O⁻ and OH⁻ anions stored in a multipole radio frequency trap have been measured using a novel laser depletion tomography method. For OH⁻ the total cross sections of $8.5(1)_{stal}(3)_{syst}$ and $8.1(1)_{stal}(7)_{syst} \times 10^{-18}$ cm², measured at 662 and 632 nm, respectively, were found constant in the temperature range of 8-300 K. The O⁻ cross sections $5.9(1)_{stal}(2)_{syst} \times 10^{-18}$ cm² measured at 170 K at 662 and 532 nm, respectively, agree within error estimations with preceding experiments and increase the accuracy of the widely used calibration standard for relative photodetachment measurements of diverse atomic and molecular species. © 2009 American Institute of Physics. [DOI: 10.1063/1.3080809]



FIG. 2. Measured OH^- ion signal as a function of storage time in the trap. The upper set of points shows the signal in the absence of the photodetachment laser. The lower points show that, when the laser is switched on at 10 s, a fast additional loss channel due to photodetachment appears. The decay rates are obtained from exponential fits (solid lines).



FIG. 1 (color online). Schematic view of the 22-pole ion trap with the rf electrodes for radial and the cylindrical dc end-cap electrodes for axial confinement. The position of the photodetachment laser is scanned along the vertical axis.



FIG. 1. (Color) Histogram of the measured photodetachment rate for O^- as a function of the transverse position of the laser light in the ion trap. The graph reflects the ion density distribution in the 22-pole trap, as the ion column density is proportional to the detachment rate. The insets show two examples for individual loss rate measurements.



FIG. 2. (Color online) Measured cross section of O^- as a function of the photon energy. Our data (large full circles) is compared with the relative measurements of Refs. 13 and 25 (squares and small triangles), which were calibrated to hydrogen anion measurements. The dashed line shows the *ab initio* calculation of Ref. 15.



Fig. 3 Photodetachment cross section of OH⁻ at a trap temperature of 50 K for varying photon energy. The steps in the cross section are due to the opening of loss channels corresponding to the J = 2, 1 and 0 rotational states of the anion.

Starting with simple ...

e⁻, H⁺, H, H⁻, H₂⁺, H₂, ..., H₃⁺





time (seconds)

103

104

H²

н³

Li⁶

10⁸

neutrons

 He^4

He³

Be⁷

Li⁷

3 X 10⁸

temperature (kelvins)

10²

1 X 10⁹

protons

neutrons

.H²

3 X 10⁹

1

10²

10⁶

10⁻¹⁰ •

10⁻¹²



In a diffuse cosmic gas of primordial composition, molecular hydrogen (H2) forms via a sequence of reactions



 H_2 molecules so formed induce the initial cooling and collapse of primordial clouds.



Doc. M. Čižek

Gerlich, D; Jusko, P; Roucka, S; Zymak, I; Plasil, R; <u>Glosik, J</u> Ion Trap Studies of $H- + H \rightarrow H2 + e-$ Between 10 and 135 K, Astrophys. J., 794 (1): , 2012.

Experimental Results for H_2 \lor Formation from H^- and H and Implications for First Star Formation

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$\mathrm{H}^- + \mathrm{H} \rightarrow \mathrm{H}_2^- \rightarrow \mathrm{H}_2 + \mathrm{e}^-$

Fig. 1. Schematic of the merged-beams apparatus used to measure the H_2 associative detachment reaction. Infrared laser photons are denoted by v_{IR} .

www.sciencemag.org SCIENCE VOL 329 2 JULY 2010

During the epoch of first star formation, molecular hydrogen (H₂) generated via associative detachment (AD) of H⁻ and H is believed to have been the main coolant of primordial gas for temperatures below 10^4 kelvin. The uncertainty in the cross section for this reaction has limited our understanding of protogalaxy formation during this epoch and of the characteristic masses and cooling times for the first stars. We report precise energy-resolved measurements of the AD reaction, made with the use of a specially constructed merged-beams apparatus. Our results agreed well with the most recent theoretically calculated cross section, which we then used in cosmological simulations to demonstrate how the reduced AD uncertainty improves constraints of the predicted masses for Population III stars.





THE ASTROPHYSICAL JOURNAL, 747:1 (6pp), 2012 ??? © 2012. The American Astronomical Society. All rights reserved. Printed in the U.S.A.

doi:10.1088/0004-637X/747/1/1

2

1

 10^{1}

10²

 T_t (K)

ION TRAP STUDIES OF $H^- + H \rightarrow H_2 + e^-$ BETWEEN 10 AND 135 K

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PHYSICAL REVIEW A 86, 032714 (2012)

Isotope effect for associative detachment: $H(D)^- + H(D) \rightarrow H_2(D_2) + e^-$

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Ion trap study of $O^- + H_2$ at low temperatures

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Determining the energy distribution of electrons produced in associative detachment: The electron spectrometer with multipole trap

P. Jusko, Š. Roučka*, R. Plašil, J. Glosík

International Journal of Mass Spectrometry 352 (2013) 19-28





Fig. 1. Schematic diagram of the ES-MPT apparatus (not to scale). The principle of combination of the radiofrequency ion trap (octopole on the left side) and the MAC-E filter (right side). Anions are injected into the octopole trap from the left and confined axially using the electrostatic ring electrodes (R₁, R₂). Laser radiation can be applied along the axis, reactant gas can be introduced into the trap between R₁ and R₂. Produced electrons are magnetically guided to the region with lower magnetic field (*B*₂) and refocused to the MCP detector. Cylindrical electrodes ESO–ES4 create the electrostatic barrier (diameter and length in parentheses are in mm).



Fig. 2. Configuration of the electrodes, coils, and ferromagnetic elements of the electron spectrometer. The calculated color-coded magnetic field intensity is projected on the calculated magnetic flux tube emanating from the ion trap. The flux tube approximately represents the possible electron trajectories. The number of Ampere-turns (At) used for producing the magnetic field is indicated for each coil. (For interpretation of colour in the artwork, the reader is referred to the web version of the article.)

ES-MPT



Storage Ion Source

αMS

Octopole

Trapping area

MAC-E filter

MCP

$O^- + H_2 \Longrightarrow H_2O + e^-$



nhotodetachment

Experimental techniques – ion traps

22PT – 22-Pole Trap

- Measurement of reaction rate coefficients
- Temperature range 10–300 K



ES-MPT – Electron Spectrometer with MultiPole Trap

- Energy distribution of detached electrons
- Octopole
- Temperature 300 K




ES-MPT









Fig. 3. Simulated spectra of monoenergetic isotropically distributed electrons with energies corresponding to the photodetached electrons from O⁻ using 660 nm (R), 532 nm (G), and 405 nm (V) lasers (see Table 1). (Panel a) Integral spectra of electrons normalized to unity at zero retarding barrier. (Panel b) Differential spectra obtained by differentiation of the spectra in panel (a). The spectra calculated without rf field (full lines), with ideal symmetric rf field (dashed lines), and with non-ideal asymmetric rf field (dotted lines) are shown.



Fig. 4. The measured normalized spectra of electrons produced by thermionic emission from hot filament. (Panel a) Integral recorded spectra. (Panels b and c) Differential spectra (normalized) obtained from the integral spectra in logarithmic and linear scales, respectively. Cathode bias voltages -0.1, -1.1, -2.1, -3.1 and -4.1 V were used to provide electrons of 5 different energy distributions (full lines). Fits of the theoretical spectrum (13) convolved with a Gaussian function and offset by the bias voltage are indicated by the dotted lines.



Fig. 8. Measured spectra of electrons photodetached from O⁻ anions. Photon wavelengths used are 660 nm (R), 532 nm (G), 445 nm (B) and 405 nm (V) (see Table 1). (Panel a) Points correspond to data measured with red, green, blue, and violet wavelengths, respectively (from left). Triangles indicate the expected electron energies according to Table 1. The smooth full lines represent the fitted convolution of the theoretical spectrum with the instrument function (Eq. (15)). The extraction probability $P_{\rm E}(U_{\rm T})$ with the trapping potential $U_{\rm T}$ determined from the fit is indicated by the dotted line and its magnitude is shown on the right axis. (Panel b) The spectra obtained from the measured data after deconvolution with the instrument function. The spectrum of the red laser cannot be deconvolved because the signal is very reduced due to the low extraction probability $P_{\rm E}$ at electron energies $\leq 0.5 \, {\rm eV}$. (For interpretation of colour in the artwork, the reader is referred to the web version of the article.)

Spectra of detached electrons (ES-MPT instrument)

$$O^- + H_2 \xrightarrow{k_1} H_2O(E_i) + e^-(E_k)$$

- Momentum conservation, Franck-Condon
- Produced molecule is highly excited





Published in [Jusko, Roučka, Plašil, Glosík; IJMS,







$$\rightarrow OH^{-} + H + 0.3 \text{ eV}$$

 $O^{-} + H_2 - H_2O + e + 3.6 eV$

Thanks

name





reaction path (a₀)

Interaction of O⁻ and H₂ at low temperatures

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(2)

$$O^- + H_2 \xrightarrow{k_1} H_2 O + e^-, \qquad \Delta H = -3.58 \text{ eV}$$
 (1)
rogen atom transfer

$$O^- + H_2 \xrightarrow{\kappa_2} OH^- + H, \qquad \Delta H = -0.28 \text{ eV},$$



FIG. 3. Panel (a)—PES of H₂O⁻ and H₂O along the minimum energy path going from O⁻ + H₂ to OH⁻ + H on the 1²A' PES. The anionic curves in the autodetachment region, where they are above the neutral PES, are indicated by points. The local minimum of the 1²A' PES, where some metastable H₂O⁻ states may exist, is magnified in the inset. In Panel (b), the path is constrained to the linear geometry, $\theta = 0^\circ$. In this case, the potential energy of the neutral H₂O is too high—outside of the graph.



FIG. 4. Panel (a)—Section of the three potential energy surfaces for $\theta = 0^{\circ}$, $r = 1.4014a_0$, showing the attraction of O⁻ and H₂ and the conical intersection at 4.6 a_0 , coupling them. The sum of quadrupole and polarization potential is marked with a dotted line. Panel (b)—Typical classical trajectory on the 2²A' PES for a collision energy of 5 meV projected on the $\theta = 0^{\circ}$ plane. The conical intersection at $\theta = 0^{\circ}$ is marked by the dashed line. The equipotential lines for V = 5 meV are also shown for the indicated values of θ . This picture shows a section of a trajectory with a duration of 700 vibrational periods of H₂. Most trajectories remain trapped for typically 10^4 – 10^5 vibrational periods, passing beyond the conical intersection several hundred times.



FIG. 5. Comparison of the measured total reaction rate coefficients $k_1 + k_2$ with the values calculated with the 1D capture model described in the text. The data are normalized with the Langevin capture rate coefficient. Sensitivity test of the model is indicated with the gray area. The part of the theoretical curve which is significantly sensitive to the parameters of the absorption potential, is shown as a dashed line. Experimental results of McFarland *et al.*¹⁰ and Viggiano *et al.*¹² are shown for comparison.

Table 2

Total reaction rate coefficient estimated from temporal evolutions of electron production at 300 K (see Fig. 10). Only the statistical error is shown. The results are accurate within a factor of 2 due to systematic errors.

$(10^{-10} \mathrm{cm^3 s^{-1}})$	[34] (10 ⁻¹⁰ cm ³ s ⁻¹)
5.4 ± 0.2	6.4
$\textbf{3.2}\pm\textbf{0.1}$	4.6
	$(10^{-10} \text{ cm}^3 \text{ s}^{-1})$ 5.4 ± 0.2 3.2 ± 0.1







Ion trap study of $O^- + H_2$ at low temperatures

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